Heat transfer in magnetohydrodynamic flow due to a linearly stretching sheet with induced magnetic field: Exact solution

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Abstract: - The solution for the magnetohydrodynamic flow, due to a linearly stretching sheet, has a simple form for the velocity field, with a companion simple form for the induced magnetic field. The associated thermal problem, including viscous dissipation and Joule heating, involves three temperature constituents, the solutions for two of which are obtained in terms of Kummer's function. The solution for the third temperature constituent is obtained in a convergent series form.

Key-Words: - MHD flow; induced magnetic field; linearly stretching sheet; heat transfer; Joule heating; viscous dissipation; streamwise diffusion; exact solution

1 Introduction

Navier-Stokes equations are inherently nonlinear due to the presence of convection terms. When convection is absent, the equations become linear and exact (closed form) solutions may be possible; e.g. Couette and Poiseuille flows [1].

With convection present, exact solutions to Navier-Stokes equations are rare. This is to the extent that reduction of order through self-similarity is regarded as exact solution [2,3].

Crane's simple exact solution [4] to the problem of the flow due to a linearly stretching sheet is one such rare solution. This invited researchers to apply the same form of solution to other related situations, which proved successful in cases involving surface feed [5], velocity slip [6], magnetohydrodynamic (MHD) flow (under the assumption of negligible induced magnetic field) [7], or combinations thereof [8].

The thermal problem associated with Crane's flow is a linear one. It was solved by Grupka and Bobba [9] neglecting viscous dissipation and streamwise heat diffusion. Neglecting Joule heating as well, Liu [10] solved the thermal problem associated with Andersson's MHD flow [7], in cases of linear surface temperature and linear heat flux.

It is shown here that the solution for the MHD flow with induced magnetic field has Crane's simple form for the velocity field, with a companion simple form for the induced magnetic field. Both solutions tend to the leading order solutions obtained in [11], as the magnetic Prandtl number diminishes. The associated thermal problem, including viscous dissipation and Joule heating, involves quadratic streamwise variation in temperature with three temperature constituents, the solutions for two of which are obtained in terms of Kummer's functions [12]. The solution for the third temperature constituent is obtained in convergent series form.

2 Mathematical model

electrically conducting, incompressible, An Newtonian fluid is driven by a non-conducting nonporous sheet, which is stretching linearly in the xdirection with rate ω . At the surface, we consider cases of prescribed temperature or heat flux. In the farfield, the fluid is essentially quiescent under pressure $p = p_{\infty}$ and temperature $T = T_{\infty}$, and is permeated by a stationary magnetic field of uniform strength B in the transverse y-direction. The velocity components in the (x,y) directions are (u,v)and the corresponding induced magnetic field components are (q,s). Constants are the fluid density ρ , kinematic viscosity ϑ , electric conductivity σ , magnetic permeability μ , specific heat c, and thermal conductivity k.

The problem admits the similarity transformations [13]

$$x = \sqrt{\vartheta/\omega}\chi, y = \sqrt{\vartheta/\omega}\eta, v = -\sqrt{\vartheta\omega}f(\eta), u = \sqrt{\vartheta\omega}\chi f'$$
 (1a,b,c,d)

$$s = B\sigma\mu\vartheta g(\eta), q = -B\sigma\mu\vartheta\chi g'$$
 (1e,f)

$$p = p_{\infty} - \rho \omega \vartheta [f' + \frac{1}{2} f^2(\eta) - \frac{1}{2} f^2(\infty)] - \frac{1}{2} B^2 \sigma^2 \mu \vartheta^2 \chi^2 g'^2$$
(1g)

$$T = T_{\infty} + \frac{\vartheta\omega}{c} [\theta_0(\eta) + \chi \theta_1(\eta) + \chi^2 \theta_2(\eta)]$$
(1h)

where primes denote differentiation with respect to the similarity coordinate η .

The governing fluid flow continuity and Navier-Stokes equations, electromagnetic Maxwell's and Ohm's equations, and energy equation including heat dissipation and Joule heating take the similarity form [13]

$$f''' + ff'' - f'^{2} = \beta [P_{\rm m} (gg'' - g'^{2}) + g'']$$
⁽²⁾

$$g'' = f' + P_{\rm m}(gf' - fg')$$
 (3)

$$\theta_2'' + \Pr(f\theta_2' - 2f'\theta_2) = -\Pr[\beta g''^2 + f''^2]$$
(4)

$$\theta_1'' + \Pr(f\theta_1' - f'\theta_1) = 0$$
(5)

$$\theta_0^{''} + \Pr f \theta_0^{'} = -2\theta_2 - 4\Pr f^{'2}$$
(6)

where $P_m = \sigma \mu \vartheta$ is the magnetic Prandtl number, $\beta = \sigma B^2 / \rho \omega$ is the magnetic interaction number, $Pr = \rho \vartheta c / k$ is the Prandtl number.

In the right-hand-sides of Eqs. (4) and (6), the terms involving f''^2 and f'^2 are due to viscous

dissipation, the term involving $g^{''2}$ is due to Joule heating, and the term involving θ_2 is due to the streamwise heat diffusion.

The boundary conditions, in the absence of velocity and thermal slip, are [13]

$$f(0) = 0, f'(0) = 1, f'(\infty) = 0$$
⁽⁷⁾

$$g(\infty) = 0, g'(\infty) = 0 \tag{8}$$

$$\theta_2(0) = \Theta_2 \text{ or } \theta_2'(0) = -Q_2, \theta_2(\infty) = 0$$
(9)

$$\theta_1(0) = \Theta_1 \text{ or } \theta_1'(0) = -Q_1, \theta_1(\infty) = 0$$
 (10)

$$\theta_0(0) = \Theta_0 \text{ or } \theta_0'(0) = -Q_0, \theta_0(\infty) = 0$$
 (11)

where Θ_0 , Θ_1 and Θ_2 are prescribed constituents of the surface temperature, while Q_0 , Q_1 and Q_2 are prescribed constituents of the heat flux from the surface.

Conditions (8) on g translate the physical requirement of the absence of the current density in the farfield, and indicate that B stands for the

farfield total magnetic field imposed and induced [11].

3 Exact solution for the velocity and magnetic fields

We attempt solutions of the form

$$f = (1 - e^{-\gamma \eta})/\gamma \tag{12}$$

$$g = \lambda e^{-\gamma \eta} \tag{13}$$

which satisfy the boundary conditions, provided $\gamma > 0$.

Substitution in Eqs. (2) and (3) gives the following relations between γ and λ .

$$\gamma^2 - 1 = \beta \gamma^2 \lambda \tag{14}$$

$$\gamma^2 \lambda = 1 + P_m \lambda \tag{15}$$

Elimination of λ results in

$$\gamma^4 - (\beta + 1 + P_m)\gamma^2 + P_m = 0$$
(16)

with the solutions

$$\gamma^{2} = \frac{1}{2} [(\beta + 1 + P_{\rm m}) \pm \sqrt{(\beta + 1 + P_{\rm m})^{2} - 4P_{\rm m}}]$$
(17)

The plus sign is chosen, so that $\gamma^2 \neq 0$ when $P_m = 0$, leading to

$$\gamma = \sqrt{\frac{1}{2} [(\beta + 1 + P_{\rm m}) + \sqrt{(\beta + 1 + P_{\rm m})^2 - 4P_{\rm m}}]}$$
(18)

and, consequently,

$$\lambda = 2[(\beta + 1 - P_m) + \sqrt{(\beta + 1 + P_m)^2 - 4P_m}]^{-1}$$
(19)

Note that both γ and λ are real, since the discriminant $(\beta + 1 + P_m)^2 - 4P_m$ manipulates to $(\beta + 1 - P_m)^2 + 4\beta P_m > 0$. Moreover, $\gamma > 1$ and $\lambda > 0$.

The solution for f reduces to Andersson's solution [7] in the case of negligible induced magnetic field ($P_m = 0$), and to Crane's solution [4] in the hydrodynamic case ($\beta = P_m = 0$). In either case, Eq. (3) merely defines g'' = f'.

4 Exact solutions for the temperature constituents

Upon substitution for f and g, Eqs. (4)-(6) combine as

$$\phi_{n}^{''} + r \left(1 - e^{-\zeta}\right) \phi_{n}^{'} - nr e^{-\zeta} \phi_{n} = \Lambda_{n}$$
(20)

where $r = \Pr/\gamma^2$, $\zeta = \gamma \eta$, $\phi_n(\zeta) = \theta_n(\eta)/\gamma^n$, and n = 2, 1 or 0, with

$$\Lambda_2 = -r(1 + \beta \gamma^2 \lambda^2) e^{-2\zeta}$$
(21)

$$= 0$$
 (22)

$$\Lambda_0 = -2\phi_2 - 4re^{-2\zeta} \tag{23}$$

and a prime denoting differentiation with respect to ζ . Henceforth, it is assumed that r > 0 is not an integer.

 Λ_1

It is realized that Eq. (20) can be transformed into Kummer's equation [12] with a source term. The solution can, thus, be expressed as the sum of a particular part due to the source term and a homogeneous part in terms of Kummer's function

$$M(a, b, z) = \sum_{m=0}^{\infty} \frac{a^{(m)}}{m! \, b^{(m)}} z^m$$

where $a^{(m)} = a(a+1) \dots (a+m-1)$ with $a^{(0)} = 1$.

In particular, the solution for ϕ_2 takes the form

$$\phi_2(\zeta) = c_0 + c_1 e^{-\zeta} + A_2 M \left(-2, 1 - r, -r e^{-\zeta}\right) + B_2 e^{-r\zeta} M \left(-2 + r, 1 + r, -r e^{-\zeta}\right)$$
(24)

where A_2 and B_2 are arbitrary constants of integration, while c_0 and c_1 are undetermined coefficients of the particular solution.

For the satisfaction of the farfield condition (9c), $A_2 = -c_0$. On the other hand, the surface condition (9a) gives

$$\Theta_2/\gamma^2 = c_0[1 - M(-2, 1 - r, -r)] + c_1 + B_2M(-2 + r, 1 + r, -r)$$
(25a)

while condition (9b) gives

$$Q_{2}/\gamma^{2} = \gamma c_{1} - \gamma c_{0} \frac{2r}{1-r} M(-1,2-r,-r) + B_{2}\gamma r[M(-2+r,1+r,-r) + \frac{2-r}{1+r} M(-1+r,2+r,-r)]$$
(25b)

Either Eq. (25a) or Eq. (25b) can be solved for B_2 in terms of c_0 and c_1 . These two coefficients are determined by the satisfaction of Eq. (20) for n=2, which gives

$$c_1 = 1 + \beta \gamma^2 \lambda^2$$
 and $c_0 = \frac{1-r}{2r} c_1$ (26)

The solution for ϕ_1 is

$$\phi_1(\zeta) = B_1 e^{-r\zeta} M(-1+r, 1 + r, -re^{-\zeta})$$
(27)

where B_1 satisfies

$$\Theta_1 / \gamma = B_1 M (-1 + r, 1 + r, -r)$$
(28a)

or

$$Q_1/\gamma = B_1 \gamma r [M(-1+r, 1+r, -r) + \frac{1-r}{1+r} M(r, 2+r, -r)]$$
(28b)

The solution for ϕ_0 is

$$\phi_0(\zeta) = B_0 e^{-r\zeta} M(r, 1+r, -re^{-\zeta}) + \varphi_0(\zeta)$$
(29)

where the particular solution φ_0 is given by

$$\varphi_{0} = -\left[4c_{0}\frac{(-r)^{2}}{(-r)^{(2)}} + 2c_{1}\right]\sum_{k=1}^{\infty}\frac{(-r)^{k}}{k(-r)^{(k+1)}}e^{-k\zeta} +\left[2c_{0}\frac{(-r)^{3}}{(-r)^{(3)}} - 4r\right]\sum_{k=2}^{\infty}\frac{(-r)^{k-2}}{k(2-r)^{(k-1)}}e^{-k\zeta} +2B_{2}e^{-r\zeta}\sum_{k=1}^{\infty}\frac{(-r)^{k}}{k!(r+k)}e^{-k\zeta}\sum_{m=1}^{k}\frac{r(3r+3m-7)+(m-2)^{2}}{(r+m)(r+m-1)(r+m-2)} +2B_{2}\zeta e^{-rx}\sum_{k=0}^{\infty}\frac{(-r)^{k}}{k!(r+k)}e^{-k\zeta}$$
(30)

and B_0 satisfies

$$\Theta_0 = B_0 M(r, 1 + r, -r) + \varphi_0(0)$$
(31a)

$$Q_0 = B_0 \gamma r \left[M(r, 1+r, -r) - \frac{r}{1+r} M(1+r, 2+r, -r) \right] - \gamma \varphi'_0(0)$$
(31b)

or

The convergence of the summations in Eq. (30) is established in Appendix A.

It is noted that, when the Θ_n 's are given, the Q_n 's in Eqs. (25b), (28b) and (31b) determine the constituents of the rate of heat transfer from the surface to the fluid. Conversely, when the Q_n 's are given, the Θ_n 's in Eqs. (25a), (28a) and (31a) determine the constituents of the surface temperature. Sample results are given in Appendix B.

5 Conclusion

It is shown that Crane's simple form of exact solution, for the hydrodynamic problem of the flow due to a linearly stretching sheet, extends to the current MHD problem with a companion simple form for the induced magnetic field. Moreover, an exact solution for the energy equation including viscous dissipation, Joule heating, and streamwise heat diffusion is obtained. The obtained exact solutions tend regularly to previously published solutions of degenerate problems [4,7,9,10].

The inclusion of the traditionally ignored physical processes of magnetic induction, viscous dissipation and Joule heating cause streamwise variations in the pressure and temperature. As Eq. (1g) indicates, the induced magnetic field instigates a favorable pressure gradient proportional to the streamwise coordinate χ . Even when the surface temperature or heat flux is streamwise uniform, viscous dissipation and (or) Joule heating bring(s) about temperature variations proportional to χ^2 .

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Appendix A

For the summation

$$S = \sum_{k=1}^{\infty} \frac{(-r)^k}{k! (r+k)} e^{-k\zeta} \sum_{m=1}^k \frac{r(3r+3m-7) + (m-2)^2}{(r+m)(r+m-1)(r+m-2)}$$

which appears in Eq. (30), we note that

$$b_m = \frac{r(3r+3m-7) + (m-2)^2}{(r+m)(r+m-1)(r+m-2)}$$

$$b_m - b_{m+1} = \frac{(m-2)(m-3) + 6r^2 + 4(m-3)r}{(r+m+1)(r+m)(r+m-1)(r+m-2)}$$

indicate that, for integer $m \ge 3$

 $b_m > b_{m+1} > 0$

regardless of the value of r. Moreover,

$$\lim_{m\to\infty}b_m=0$$

With

$$C_k = \sum_{m=3}^k b_m > 0$$

we have

$$\lim_{k \to \infty} \frac{C_{k+1}}{C_k} = \lim_{k \to \infty} (1 + \frac{b_{k+1}}{C_k}) = 1$$

and we can write

$$S = -\frac{(3r-1)}{(r+1)^2}e^{-\zeta} + \frac{r(3r-1)}{(r+2)^2}e^{-2\zeta} + \frac{2(3r-1)}{(r+2)r}\sum_{k=3}^{\infty}\frac{(-r)^k}{k!(r+k)}e^{-k\zeta} + \sum_{k=3}^{\infty}\frac{(-r)^k}{k!(r+k)}C_ke^{-k\zeta}$$

Thus, the infinite summations included in Eq. (30) have the forms

$$\sum_{k} \frac{(-r)^{k}}{k(-r)^{(k+1)}} e^{-k\zeta}, \sum_{k} \frac{(-r)^{k-2}}{k(2-r)^{(k-1)}} e^{-k\zeta}, \sum_{k} \frac{(-r)^{k}}{k! (r+k)} e^{-k\zeta}, \sum_{k} \frac{(-r)^{k}}{k! (r+k)} C_{k} e^{-k\zeta}$$

whose convergence can be easily established through the ratio test.

Appendix B

The exact solutions for the thermal problem are used to produce sample numerical values, in order to demonstrate their usefulness.

In Table 1, we compare the present results to previously published results. The problem under consideration is one with linearly varying surface temperature ($\Theta_0 = 0$, $\Theta_1 = 1$, $\Theta_2 = 0$). The physical processes of viscous dissipation and streamwise diffusion are ignored in [9] and in [10], which ignores magnetic induction and Joule heating, as well. Only one heat transfer constituent, Q_1 , is involved. Retaining the ignored processes, revives the other two constituents Q_0 and Q_2 . Further, comparison with the numerically obtained results in [13] indicates the high accuracy of the numerical method adopted in [13]. This illustrates one benefit of exact solutions.

Case	Ref.	Q_0	Q_1	Q_2	Q_0	Q_1	Q_2
$\beta = 0, P_{\rm m} = 0$ $Pr = 0.72$	[9]	0	0.8086	0	-1.49116	0.80863	-0.25635
$\beta = 1, P_{\rm m} = 0$ $Pr = 0.7$	[10]	0	0.689699	0	-1.52369	0.689699	-0.59691
$\beta = 1, P_{\rm m} = 0.1$ $Pr = 0.72$	[13]	-1.56122	0.69997	-0.63757	-1.56119	0.69997	-0.63758
						Present	

Table 1 Comparison of present results to literature results.

The effect of neglecting one or all of the abovementioned thermal processes is presented in Tables 2a,b for Case 1 when the surface is maintained at the freestream temperature, and in

Tables 3a,b for Case 2 when the surface is thermally insulated. In either case, θ_1 is governed by a homogeneous problem, obviating its contribution. Moreover, profiles for the temperature gradients θ'_0 and θ'_2 in Case 1, and for the temperature constituents θ_0 and θ_2 in Case 2 are shown in Figs. 1 and 2, respectively.

The following is reported.

- The contribution of the induced magnetic field is minor, even at the high value of the magnetic Prandtl number $P_m = 0.1$, invoked throughout.
- Heat generated within the fluid, through viscous friction and Joule heating, results in

heat flux to the surface in Case 1, and in rise in surface temperature in Case 2.

- The streamwise diffusion, which is ignored in boundary layer formulations, is of appreciable effect, increasing as the magnetic interaction number β increases and as the Prandtl number Pr decreases.
- Contributions of the constituent θ_2 magnify downstream, being multiplied by χ^2 .

Results for Case 1: Constant Surface Temperature

Table 2 Effect of magnetic induction, viscous dissipation, Joule heating and streamwise diffusion. a. Pr = 0.72

	Neglecting	Neglecting	Neglecting	Neglecting	Neglecting	Retaining	
Case	All	Magnetic	Viscous	Joule	Streamwise	All	β
	Effects	Induction	Dissipation	Heating	Diffusion	Effects	
Q_0	0	-1.48466	-0.02879	-1.45805	-1.19353	-1.48684	0.1
Q_2	0	-0.29860	-0.02997	-0.27540	-0.30537		0.1
Q_0	0	-1.54337	-0.22394	-1.33725	-0.92158	-1.56119	1
Q_2	0	-0.61143	-0.22323	-0.41435	-0.63758		1
Q_0	0	-3.09839	-1.29233	-1.83132	-0.42373	-3.12365	10
Q_2	0	-2.15077	-1.03905	-1.13172	-2.17077		10
b. Pr =	7.0						
	Neglecting	Neglecting	Neglecting	Neglecting	Neglecting	Retaining	
Case	All	Magnetic	Viscous	Joule	Streamwise	All	β
	Effects	Induction	Dissipation	Heating	Diffusion	Effects	
Q_0	0	-8.71366	-0.02238	-8.67228	-8.4666	-8.69466	0.1
Q_2	0	-1.62865	-0.16367	-1.50397	-1.66764		0.1
Q_0	0	-7.42052	-0.15854	-7.22010	-6.92583	-7.37864	1
Q_2	0	-3.62323	-1.32751	-2.46403	-3.79155		1
Q_0	0	-5.28842	-0.77206	-4.52084	-3.67993	-5.29290	10
Q_2	0	-16.14890	-7.81008	-8.50665	-16.31	-16.31674	



Fig. 1 Temperature Gradients. $\beta = 1$, $P_m = 0.1$, Pr = 0.72.

Results for Case 2: Thermally Insulated Surface	
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Table 3 Effect of magnetic induction, viscous dissipation, Joule heating and streamwise diffusion. a. Pr = 0.72

Naglasting	Naglasting	Naglasting	Noglasting	Naglasting	Dotoining	
Neglecting	Neglecting	Neglecting	Neglecting	Regietting	Ketanning	0
All	Magnetic	Viscous	Joule	Streamwise	All	β
Effects	Induction	Dissipation	Heating	Diffusion	Effects	
0	4.38432	0.17596	4.25335	2.63645	4.42930	0.1
0	0.27786	0.0 2792	0.25659	0.28452		0.1
0	8.02607	2.10342	6.29289	2.38869	8.39631	1
0	0.62938	0.23097	0.42871	0.65967		1
0	157.40217	75.90184	84.75464	2.08323	160.65647	10
0	3.71306	1.79966	1.96016	3.75982		10
r = 7.0						
Neglecting	Neglecting	Neglecting	Neglecting	Neglecting	Retaining	
All	Magnetic	Viscous	Joule	Streamwise	All	β
Effects	Induction	Dissipation	Heating	Diffusion	Effects	
0	4.69997	0.01960	4.67477	4.49464	4.69437	0.1
0	0.41139	0.04135	0.38001	0.42136		0.1
0	4.30634	0.15985	4.14265	3.84595	4.30251	1
0	0.93778	0.34401	0.63853	0.98255		1
0	5.79220	1.53860	4.29908	2.62325	5.83768	10
0	4.83854	2.34280	2.55175	4.89454		10
	$\begin{tabular}{c} Neglecting \\ All \\ Effects \\\hline 0 \\ 0 \\\hline 0 \\\hline 0 \\\hline 0 \\\hline 0 \\\hline 0 \\\hline 0$	Neglecting AllNeglecting MagneticEffectsInduction0 4.38432 0 0.27786 0 8.02607 0 0.62938 0 157.40217 0 3.71306 r = 7.0Neglecting MagneticEffectsInduction0 4.69997 0 0.41139 0 4.30634 0 0.93778 0 5.79220 0 4.83854	Neglecting AllNeglecting MagneticNeglecting ViscousEffectsInductionDissipation0 4.38432 0.17596 0 0.27786 0.02792 0 8.02607 2.10342 0 0.62938 0.23097 0 157.40217 75.90184 0 3.71306 1.79966 r = 7.0Neglecting MagneticNeglecting ViscousEffectsInductionDissipation0 4.69997 0.01960 0 0.41139 0.04135 0 4.30634 0.15985 0 0.93778 0.34401 0 5.79220 1.53860 0 4.83854 2.34280	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $	$\begin{array}{c c c c c c c c c c c c c c c c c c c $



Fig. 2 Temperature Constituents. $\beta = 1$, $P_m = 0.1$, Pr = 0.72.