# New First Integrals for the Continuity, Vorticity and Related Equations

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Abstract: This work studies and clarifies some local phenomena in fluid mechanics, in the form of an intrinsic analytic study, regarding the continuity equation, its first integral (the flow rate equation), for inviscid compressible fluids, and the vorticity equation, for viscous incompressible fluids, finding new first integrals. It continues a series of works presented at some conferences and a congress during 2006 - 2012, representing a real deep insight into the still hidden theory of the isoenergetic rotational flow. Several new functions, surfaces and vectors were introduced: the 2-D "quasi-stream" function on the 3-D (V,  $\Omega$ ) surfaces, for the continuity equation; the surfaces of *iso*-normal mass flux *density* (over which the continuity equation for the steady flow in a thick stream tube admits the same first integral as for the flow in a thin one, and whose envelope sheets are the sections of uniform flow), for the flow rate equation; the 3-D stream function vector, allowing new local and global forms of continuity equation (the global one similar to Helmholtz' 2nd theorem about vortices in an ideal fluid); Selescu's incompressible roto-viscous vector and the zero-work surfaces (for some non-conservative vectors), for the vorticity equation. The dependence "2-D velocity quasi-potential  $\leftrightarrow$  2-D quasi-stream function" was established. The unsteady flow's continuity equation was analytically integrated.

*Key-Words:* rotational flows; steady and unsteady flows; inviscid and viscous fluids; compressible fluids; isentropic or  $(V, \Omega)$  surfaces; Selescu's 3-D stream function vector, surfaces of *iso*-normal mass flux *density*, roto-viscous vector and zero-work surfaces

# 1 Introduction, nomenclature and the first approach to the new steady model

The continuity equation is the mathematical expression of the law of conservation of mass in fluid mechanics. For a certain 3-D unsteady motion, if there are no sources or losses of fluid within the considered flow subdomain, the general form of this equation (called its local form) is:

$$\frac{\partial \rho}{\partial t} + \nabla(\rho \mathbf{V}) = 0 \quad \text{(Euler), or } \frac{d\rho}{dt} + \rho \nabla \mathbf{V} = 0 \text{, where}$$
$$\nabla = \sum_{i=1}^{3} \frac{\mathbf{k}_{i}}{\mathbf{h}_{i}} \frac{\partial}{\partial x_{i}} = \mathbf{k}_{x} \frac{\partial}{\partial x} + \mathbf{k}_{y} \frac{\partial}{\partial y} + \mathbf{k}_{z} \frac{\partial}{\partial z} - \text{nabla}$$

(Hamilton's operator), in a triorthogonal system of curvilinear coordinates  $\mathbf{x}_i$ ;  $\mathbf{k}_i - a$  3-D basis;  $\mathbf{h}_i - \text{Lamé's coefficients}$ ;  $\mathbf{V}$  – the local instantaneous velocity of translation (of a small fluid particle) – the intensity of the local fluid field;  $\mathbf{V} = |\mathbf{V}|$ ;  $\mathbf{\Omega} = \nabla \times \mathbf{V} = 2\boldsymbol{\omega}$  – the vorticity, with:  $\boldsymbol{\omega}$  – the local instantaneous velocity of rotation (of the particle); Oxyz – the Cartesian orthogonal coordinate system;  $\rho$  – the fluid local density; t – the time. Additionally: a – the speed of sound; c – the critical speed (V = a). For a certain 3-D steady motion, studied in a certain triorthogonal curvilinear  $q_i$  (i = 1, 3) coordinate system (like the previously used Oξηζ or Oλµv – both intrinsic, with:  $d\lambda = h_\xi d\xi$ ;  $d\mu = h_\eta d\eta$  and  $dv = h_\zeta d\zeta$  – the elementary arc lengths), the continuity equation yields the forms:  $\nabla(\rho \mathbf{V}) = 0$  (with  $\mathbf{V} = \mathbf{k}_{q_1} V_{q_1} + \mathbf{k}_{q_2} V_{q_2} + \mathbf{k}_{q_3} V_{q_3}$ ), or:

$$\frac{1}{\prod_{j=1}^{3} h_{j}} \sum_{i=1}^{3} \frac{\partial}{\partial q_{i}} \left[ \rho \left( \prod_{j=1}^{3} h_{j} \right) \frac{V_{q_{i}}}{h_{i}} \right] = 0 \quad , \quad \text{or more}$$

$$\frac{1}{\prod_{j=1}^{3} h_{j}} \sum_{i=1}^{3} \frac{\partial}{\partial q_{i}} \left[ \rho \left( \prod_{j=1}^{3} h_{j} \right) \dot{q}_{i} \right] = 0 \quad , \quad (1)$$

meaning the mass flux density  $\rho \mathbf{V}$  is a solenoidal vector field ( $\rho \mathbf{V} = \nabla \times \Psi_c$  with  $\Psi_c \neq \nabla G$ ; if  $\Psi_c = \nabla G$  one can define the flow's stagnation lines – see next section);  $h_i$  are the corresponding Lamé's coefficients ( $h_{\xi}$ ,  $h_{\eta}$ ,  $h_{\zeta}$ ):

$$\mathbf{h}_{i} = \sqrt{\sum_{j=1}^{3} \left(\frac{\partial \mathbf{x}_{j}}{\partial \mathbf{q}_{i}}\right)^{2}} \quad ; \quad (i = \overline{1, 3}) \quad ; \quad \text{with}:$$

 $\int x_j = x, y, z - Cartesian coordinates;$ 

 $q_i = \xi, \eta, \zeta$  – orthogonal curvilinear coordinates. Eq. (1) may be written explicitly in the forms below:

$$\begin{split} & \frac{1}{h_{\xi}h_{\eta}h_{\zeta}} \Bigg[ \frac{\partial}{\partial\xi} \Bigg( \rho h_{\xi}h_{\eta}h_{\zeta} \frac{V_{\xi}}{h_{\xi}} \Bigg) + \frac{\partial}{\partial\eta} \Bigg( \rho h_{\xi}h_{\eta}h_{\zeta} \frac{V_{\eta}}{h_{\eta}} \Bigg) \\ & + \frac{\partial}{\partial\zeta} \Bigg( \rho h_{\xi}h_{\eta}h_{\zeta} \frac{V_{\zeta}}{h_{\zeta}} \Bigg) \Bigg] = 0 \quad , \quad \text{or} \\ & \frac{1}{h_{\xi}h_{\eta}h_{\zeta}} \Bigg[ \frac{\partial}{\partial\xi} \Bigl( \rho h_{\eta}h_{\zeta}V_{\xi} \Bigr) + \frac{\partial}{\partial\eta} \Bigl( \rho h_{\xi}h_{\zeta}V_{\eta} \Bigr) + \frac{\partial}{\partial\zeta} \Bigl( \rho h_{\xi}h_{\eta}V_{\zeta} \Bigr) \Bigg] \\ & = 0 \quad , \quad \text{or with:} V_{\xi} = h_{\xi}\dot{\xi} \quad ; \quad V_{\eta} = h_{\eta}\dot{\eta} \quad ; \quad V_{\zeta} = h_{\zeta}\dot{\zeta} \quad ; \end{split}$$

$$\frac{1}{h_{\xi}h_{\eta}h_{\zeta}}\left[\frac{\partial}{\partial\xi}\left(\rho h_{\xi}h_{\eta}h_{\zeta}\dot{\xi}\right)+\frac{\partial}{\partial\eta}\left(\rho h_{\xi}h_{\eta}h_{\zeta}\dot{\eta}\right)+\frac{\partial}{\partial\zeta}\left(\rho h_{\xi}h_{\eta}h_{\zeta}\dot{\zeta}\right)\right]$$
  
=0, or more  
$$\frac{1}{h_{\xi}h_{\eta}h_{\zeta}}\left[\left(\dot{\xi}\frac{\partial}{\partial\xi}+\dot{\eta}\frac{\partial}{\partial\eta}+\dot{\zeta}\frac{\partial}{\partial\zeta}\right)\left(\rho h_{\xi}h_{\eta}h_{\zeta}\right)\right]$$
+ $\rho h_{\xi}h_{\eta}h_{\zeta}\left(\frac{\partial\dot{\xi}}{\partial\xi}+\frac{\partial\dot{\eta}}{\partial\eta}+\frac{\partial\dot{\zeta}}{\partial\zeta}\right)\right]$ =0;  
$$\rho\left[\frac{1}{\rho h_{\xi}h_{\eta}h_{\zeta}}\left(\frac{d\xi}{dt}\frac{\partial}{\partial\xi}+\frac{d\eta}{dt}\frac{\partial}{\partial\eta}+\frac{d\zeta}{dt}\frac{\partial}{\partial\zeta}\right)\left(\rho h_{\xi}h_{\eta}h_{\zeta}\right)\right]$$
+ $\left(\frac{\partial\dot{\xi}}{\partial\xi}+\frac{\partial\dot{\eta}}{\partial\eta}+\frac{\partial\dot{\zeta}}{\partial\zeta}\right)$ =0, ( $\rho \neq 0$ ), this resulting in  
$$\frac{1}{\rho h_{\xi}h_{\eta}h_{\zeta}}\frac{d}{dt}\left(\rho h_{\xi}h_{\eta}h_{\zeta}\right)+\frac{\partial\dot{\xi}}{\partial\xi}+\frac{\partial\dot{\eta}}{\partial\eta}+\frac{\partial\dot{\zeta}}{\partial\zeta}=0, \text{ or more}$$
$$\left[\ln\left(\rho h_{\xi}h_{\eta}h_{\zeta}\right)\right]+\frac{\partial\dot{\xi}}{\partial\xi}+\frac{\partial\dot{\eta}}{\partial\eta}+\frac{\partial\dot{\zeta}}{\partial\zeta}=0, \quad (2)$$

or, in an expanded form

$$\frac{\dot{\rho}}{\rho} + \frac{\dot{h}_{\xi}}{h_{\xi}} + \frac{\dot{h}_{\eta}}{h_{\eta}} + \frac{\dot{h}_{\zeta}}{h_{\zeta}} + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} = 0 \quad , \tag{3}$$

or, taking into account that on any "i" (**V**, **Ω**) surface  $(\zeta = \zeta_{0i})$  we have:  $V_{\zeta} = h_{\zeta}\dot{\zeta} = \dot{v} = 0$  and so  $\partial(\rho h_{\xi} h_{\eta} V_{\zeta})/\partial \zeta$ = 0, (where  $\mathbf{\Omega} = \nabla \times \mathbf{V} = 2\boldsymbol{\omega}$  is the local vorticity, curl **V**):

$$\frac{1}{\mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}} \left[ \frac{\partial}{\partial\xi} \left( \rho \mathbf{h}_{\eta} \mathbf{h}_{\zeta} \mathbf{V}_{\xi} \right) + \frac{\partial}{\partial\eta} \left( \rho \mathbf{h}_{\xi} \mathbf{h}_{\zeta} \mathbf{V}_{\eta} \right) \right] = 0 \quad , \quad \text{or} \\ \frac{1}{\mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}} \left[ \frac{\partial}{\partial\xi} \left( \rho \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta} \dot{\xi} \right) + \frac{\partial}{\partial\eta} \left( \rho \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta} \dot{\eta} \right) \right] = 0 \quad . \quad (4)$$

Let us introduce on any  $(\mathbf{V}, \boldsymbol{\Omega})$  surface  $(\zeta = \zeta_{0i})$  the 2-D velocity "quasi-potential"  $\Phi_i(\xi, \eta)$  (see [1]–[4] and section 3):

$$\begin{split} \mathbf{V}_{\boldsymbol{\xi}i} &= \mathbf{V}_{\boldsymbol{\lambda}i} = \frac{1}{h_{\boldsymbol{\xi}}} \frac{\partial \Phi_{i}}{\partial \boldsymbol{\xi}} = \frac{\partial \Phi_{i}}{\partial \boldsymbol{\lambda}}; \\ \mathbf{V}_{\boldsymbol{\eta}i} &= \mathbf{V}_{\boldsymbol{\mu}i} = \frac{1}{h_{\boldsymbol{\eta}}} \frac{\partial \Phi_{i}}{\partial \boldsymbol{\eta}} = \frac{\partial \Phi_{i}}{\partial \boldsymbol{\mu}}; \\ \mathbf{V}_{\boldsymbol{\zeta}i} &= \mathbf{V}_{\boldsymbol{\nu}i} = \frac{1}{h_{\boldsymbol{\zeta}}} \frac{\partial \Phi_{i}}{\partial \boldsymbol{\zeta}} = 0; \quad \mathbf{V}_{i} = \nabla \Phi_{i} \quad . \end{split}$$

The mass flux density  $(\rho \mathbf{V})_i = (\rho \nabla \Phi)_i$  having also two components only, Eqs. (2) and (4) become respectively:

$$\begin{split} & \left[ \ln \left( \rho h_{\xi} h_{\eta} h_{\zeta} \right) \right] + \frac{\partial \xi}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} = 0 , \quad \text{and} \\ & \frac{1}{h_{\xi} h_{\eta} h_{\zeta}} \left[ \frac{\partial}{\partial \xi} \left( \rho \frac{h_{\eta} h_{\zeta}}{h_{\xi}} \frac{\partial \Phi_{i}}{\partial \xi} \right) + \frac{\partial}{\partial \eta} \left( \rho \frac{h_{\xi} h_{\zeta}}{h_{\eta}} \frac{\partial \Phi_{i}}{\partial \eta} \right) \right] = 0 , \\ & \text{or} \quad \frac{\partial}{\partial \xi} \left( \rho \frac{h_{\eta} h_{\zeta}}{h_{\xi}} \frac{\partial \Phi_{i}}{\partial \xi} \right) = \frac{\partial}{\partial \eta} \left( - \rho \frac{h_{\xi} h_{\zeta}}{h_{\eta}} \frac{\partial \Phi_{i}}{\partial \eta} \right) . \end{split}$$

#### 2 Introducing the 2-D "quasi-stream function" on the (V, $\Omega$ ) surfaces and Selescu's 3-D "stream function" vector (both for the mass flux density vector)

The last PDE is possible if and only if there is a scalar function  $F_i(\xi, \eta, \zeta_{0i})$  (a 2-D "quasi-stream function" for the mass flux density vector  $\mathbf{j} = \rho \mathbf{V}$  on the surface  $\zeta = \zeta_{0i}$ ) so that the conditions below are fulfilled:

$$\begin{pmatrix} \rho \frac{h_{\eta}h_{\zeta}}{h_{\xi}} \frac{\partial \Phi}{\partial \xi} \end{pmatrix}_{i} = \frac{\partial F_{i}}{\partial \eta}; \quad \left( \rho \frac{h_{\xi}h_{\zeta}}{h_{\eta}} \frac{\partial \Phi}{\partial \eta} \right)_{i} = -\frac{\partial F_{i}}{\partial \xi}, \quad \text{or}: \\ \left( \rho \frac{h_{\zeta}}{h_{\xi}} \frac{\partial \Phi}{\partial \xi} \right)_{i} = \frac{\partial F_{i}}{h_{\eta} \partial \eta}; \quad \left( \rho \frac{h_{\zeta}}{h_{\eta}} \frac{\partial \Phi}{\partial \eta} \right)_{i} = -\frac{\partial F_{i}}{h_{\xi} \partial \xi} \quad ,$$

to satisfy Schwarz' theorem for the mixed derivatives of the second order  $(\partial^2 F_i/\partial \xi \partial \eta = \partial^2 F_i/\partial \eta \partial \xi)$ . So we can write:

$$dF_{i} = \left(\rho \frac{h_{\eta}h_{\zeta}}{h_{\xi}} \frac{\partial \Phi}{\partial \xi}\right)_{i} d\eta - \left(\rho \frac{h_{\xi}h_{\zeta}}{h_{\eta}} \frac{\partial \Phi}{\partial \eta}\right)_{i} d\xi \quad ,$$

a total differential admitting along flow's "quasi-streamlines"  $dF_i = 0$  ("mass flux density" streamlines) the first integral:  $F_i(\xi, \eta, \zeta_{0i}) = \text{const.}$ 

With the elementary orthogonal arcs:  $d\lambda = h_{\xi}d\xi$  and  $d\mu = h_{\eta}d\eta$ , the conditions above become:

$$\frac{\partial F_{i}(\lambda,\mu,\nu_{0i})}{\partial \lambda_{i}} = -\left(\rho h_{\zeta} \frac{\partial \Phi}{\partial \mu}\right)_{i}; \frac{\partial F_{i}(\lambda,\mu,\nu_{0i})}{\partial \mu_{i}} = \left(\rho h_{\zeta} \frac{\partial \Phi}{\partial \lambda}\right)_{i};$$

$$\frac{\partial F_{i}(\xi,\eta,\zeta_{0i})}{(h_{\xi}\partial\xi)_{i}} = -\left(\rho \frac{h_{\zeta}}{h_{\eta}} \frac{\partial \Phi}{\partial \eta}\right)_{i}; \frac{\partial F_{i}(\xi,\eta,\zeta_{0i})}{(h_{\eta}\partial\eta)_{i}} = \left(\rho \frac{h_{\zeta}}{h_{\xi}} \frac{\partial \Phi}{\partial\xi}\right)_{i},$$
or more:
$$\begin{cases} \left(\frac{\partial \Phi}{\partial\xi}\right)_{i} = \left[\frac{1}{\rho} \frac{h_{\xi}}{h_{\zeta}} \frac{\partial F(\xi,\eta,\zeta)}{h_{\eta}\partial\eta}\right]_{\zeta=\zeta_{0i}}; \\ \left(\frac{\partial \Phi}{\partial\eta}\right)_{i} = -\left[\frac{1}{\rho} \frac{h_{\eta}}{h_{\zeta}} \frac{\partial F(\xi,\eta,\zeta)}{h_{\xi}\partial\xi}\right]_{\zeta=\zeta_{0i}}. \end{cases}$$

For a compressible fluid flow in a certain triorthogonal curvilinear coordinate system  $O\xi\eta\zeta$  (of elementary orthogonal arcs  $d\lambda = h_{\xi}d\xi$ ,  $d\mu = h_{\eta}d\eta$  and  $dv = h_{\zeta}d\zeta$ ), we can introduce (by definition) a new 3-D "stream function"  $\Psi_c$  (Selescu's vector), this being given by a relation involving the analytic expression below of the mass flux density vector  $\mathbf{j} = \rho \mathbf{V}$  in the 3-D intrinsic basis  $\mathbf{k}_i$  (so being obtained *a new local form for the continuity equation*):

$$\mathbf{j} = \rho \mathbf{V} = \nabla \times \mathbf{\Psi}_{c} = \frac{1}{\mathbf{h}_{\xi} \mathbf{h}_{\eta} \mathbf{h}_{\zeta}} \begin{vmatrix} \mathbf{h}_{\xi} \mathbf{k}_{\xi} & \mathbf{h}_{\eta} \mathbf{k}_{\eta} & \mathbf{h}_{\zeta} \mathbf{k}_{\zeta} \\ \frac{\partial}{\partial \xi} & \frac{\partial}{\partial \eta} & \frac{\partial}{\partial \zeta} \\ \mathbf{h}_{\xi} \Psi_{c\xi} & \mathbf{h}_{\eta} \Psi_{c\eta} & \mathbf{h}_{\zeta} \Psi_{c\zeta} \end{vmatrix}$$
$$= \begin{vmatrix} \mathbf{k}_{\xi} & \mathbf{k}_{\eta} & \mathbf{k}_{\zeta} \\ \frac{\partial}{\partial \lambda} & \frac{\partial}{\partial \mu} & \frac{\partial}{\partial v} \\ \Psi_{c\xi} & \Psi_{c\eta} & \Psi_{c\zeta} \end{vmatrix} = \mathbf{k}_{\xi} \rho \mathbf{V}_{\xi} + \mathbf{k}_{\eta} \rho \mathbf{V}_{\eta} + \mathbf{k}_{\zeta} \rho \mathbf{V}_{\zeta} \quad ,$$

with  $\Psi_c \neq \nabla G$  (G being a certain scalar function), except for some space curves – singular lines of a saddle point type (bifurcation lines), along which  $\Psi_c = \nabla G_{ij}$ , this meaning the existence of steady flow's "ij" stagnation lines, they being 0-vorticity lines too. The subscripts "i" and "j" denote two families of *orthogonal* surfaces, by whose intersection the "ij" curves are obtained; so, along these "ij" space curves the mass flux density vector  $\mathbf{j} = \rho \mathbf{V} = 0$ . The equation  $\Psi_c = \nabla G_{ij}$  is the definition of flow's stagnation lines. If  $\Psi_c$  is just  $k\rho \mathbf{V}$ , so having  $\mathbf{j} = \rho \mathbf{V} = k\nabla \times \rho \mathbf{V} = k\nabla \times \mathbf{j}$ , one gets a special compressible helicoidal (screw) motion – a special Beltrami flow. On any "i" isentropic ( $\mathbf{V}, \Omega$ ) surface ( $\zeta = \zeta_{0i}$ , see [1]–[4]) of elementary orthogonal arcs  $d\lambda = h_{\xi}d\xi$  and  $d\mu = h_{\eta}d\eta$ , we have for this compressible flow:  $\Psi_{ci} = \mathbf{k}_{\zeta}\Psi_{ci}$  with  $\Psi_{ci} =$  $\Psi_{c\zeta}(\xi, \eta, \zeta_{0i}) = \Psi_{c\zeta}(\lambda, \mu, v_{0i})$  ( $\Psi_{c\zeta j} = \Psi_{c\eta j} = 0$ ) and respectively:

$$\begin{aligned} \mathbf{j}_{i} &= \rho \mathbf{V}_{i} = \operatorname{curl} \mathbf{\Psi}_{ci} = \nabla \times \mathbf{\Psi}_{ci} = \frac{1}{h_{\xi} h_{\eta} h_{\zeta}} \begin{vmatrix} h_{\xi} \mathbf{k}_{\xi} & h_{\eta} \mathbf{k}_{\eta} & h_{\zeta} \mathbf{k}_{\zeta} \\ \frac{\partial}{\partial \xi} & \frac{\partial}{\partial \eta} & \frac{\partial}{\partial \zeta} \\ 0 & 0 & h_{\zeta} \Psi_{ci} \end{vmatrix} \\ &= \begin{vmatrix} \mathbf{k}_{\xi} & \mathbf{k}_{\eta} & \mathbf{k}_{\zeta} \\ \frac{\partial}{\partial \lambda} & \frac{\partial}{\partial \mu} & \frac{\partial}{\partial v} \\ 0 & 0 & \Psi_{ci} \end{vmatrix} = \mathbf{k}_{\xi} j_{\xi i} + \mathbf{k}_{\eta} j_{\eta i} = \mathbf{k}_{\xi} \rho \mathbf{V}_{\xi i} + \mathbf{k}_{\eta} \rho \mathbf{V}_{\eta i} ; \Psi_{ci} \neq \nabla \mathbf{G}_{i} , \end{aligned}$$

like for a local quasi-plane 2-D fluid motion (in the  $\xi O\eta$ plane tangent to the "i" isentropic surface), this leading to:  $j_{\xi i} = \rho V_{\xi i} = \frac{1}{h_{\eta}} \frac{\partial \Psi_{\dot{\alpha}}}{\partial \eta} = \frac{\partial \Psi_{\dot{\alpha}}}{\partial \mu}; j_{\eta i} = \rho V_{\eta i} = -\frac{1}{h_{\xi}} \frac{\partial \Psi_{\dot{\alpha}}}{\partial \xi} = -\frac{\partial \Psi_{\dot{\alpha}}}{\partial \lambda};$ 

 $(j_{\zeta i} = \rho V_{\zeta i} = 0)$ , the newly introduced vector  $\Psi_c$  "becoming" now a 2-D "quasi-stream" scalar function  $\Psi_{ci}$ .

#### **3** Introducing the 2-D velocity "quasipotential" over the isentropic surfaces

The vector **V** has now two components only (like the vector  $\Omega$ , both lying in the plane tangent to an isentropic sheet). Let be Oxyz the Cartesian system. In a triorthogonal curvilinear coordinate system O $\xi\eta\zeta$  tied to this isentropic surface (having  $\xi$ O $\eta$  as tangent plane) – therefore *a smart intrinsic coordinate system*, the vorticity component normal to the isentropic sheet ( $\zeta = \zeta_0$ ) must be:  $\Omega_{\zeta} = 0$ . The analytic expression of the vector  $\Omega$  is:

$$\begin{split} \mathbf{\Omega} = \nabla \times \mathbf{V} = & \frac{1}{\mathbf{h}_{\xi} \mathbf{h}_{\eta} \mathbf{h}_{\zeta}} \begin{vmatrix} \mathbf{h}_{\xi} \mathbf{k}_{\xi} & \mathbf{h}_{\eta} \mathbf{k}_{\eta} & \mathbf{h}_{\zeta} \mathbf{k}_{\zeta} \\ \frac{\partial}{\partial \xi} & \frac{\partial}{\partial \eta} & \frac{\partial}{\partial \zeta} \end{vmatrix} = & \mathbf{k}_{\xi} \mathbf{\Omega}_{\xi} + & \mathbf{k}_{\eta} \mathbf{\Omega}_{\eta} + & \mathbf{k}_{\zeta} \mathbf{\Omega}_{\zeta}, \\ \mathbf{h}_{\xi} \mathbf{V}_{\xi} & \mathbf{h}_{\eta} \mathbf{V}_{\eta} & & \mathbf{h}_{\zeta} \mathbf{V}_{\zeta} \end{vmatrix} = & \mathbf{k}_{\xi} \mathbf{\Omega}_{\xi} + & \mathbf{k}_{\eta} \mathbf{\Omega}_{\eta} + & \mathbf{k}_{\zeta} \mathbf{\Omega}_{\zeta}, \\ \text{with} : \mathbf{V}_{\xi} = & \mathbf{h}_{\xi} \dot{\xi} ; \mathbf{V}_{\eta} = & \mathbf{h}_{\eta} \dot{\eta} ; \mathbf{V}_{\zeta} = & \mathbf{h}_{\zeta} \dot{\zeta} = & \mathbf{0} \end{split}$$

and equating to zero the expression of component  $\Omega_{\zeta}$ :

$$\Omega_{\zeta} = \frac{1}{h_{\xi}h_{\eta}} \left[ \frac{\partial(h_{\eta}V_{\eta})}{\partial\xi} - \frac{\partial(h_{\xi}V_{\xi})}{\partial\eta} \right] = 0 , \text{ or } 
\frac{\partial(h_{\eta}V_{\eta})}{\partial\xi} - \frac{\partial(h_{\xi}V_{\xi})}{\partial\eta} = 0 .$$

Between two isentropic sheets the specific entropy S can vary continuously (monotonously) or discontinuously (by jump, like for the cases of supersonic plane flow with direct and Mach reflected shocks, and of axisymmetric confluent flows). Even if varying the index "i",  $S = S_{0i}(i)$  is not always a strictly increasing function to be accepted as a  $\zeta$  coordinate (like for symmetric plane flows), the monotony of S on some intervals of "i" may be considered, thus needing delimiters. Let introduce a scalar function  $\Phi_i(M) = \Phi_i(\xi, \eta, \zeta_{0i})$ , called by the author "quasi-potential", whose partial derivatives along the directions of the elementary orthogonal arcs  $h_z d\xi$ and  $h_n d\eta$  on the "i" isentropic surface ( $\zeta = \zeta_{0i}$ ) are just the components  $V_{\xi i}$  and  $V_{\eta i}$  of the velocity  $\mathbf{V}_i$  vector ( $V_{\xi i} = 0$ ). Let be d**R** an elementary virtual displacement vector in the plane tangent to an isentropic sheet (coplanar to V and  $\Omega$ ):  $d\mathbf{R} = c_1 \mathbf{V} + c_2 \mathbf{\Omega} = \mathbf{k}_{\xi} h_{\xi} d\xi + \mathbf{k}_n h_n d\eta = \mathbf{k}_{\xi} d\lambda + \mathbf{k}_n d\mu$ , the elementary arc length ds (=  $|d\mathbf{R}|$ ) on this surface being given by ds<sup>2</sup> = dx<sup>2</sup> + dy<sup>2</sup> + dz<sup>2</sup> = Ed\xi<sup>2</sup> + 2Fd\xid\eta + Gd\eta<sup>2</sup> =  $h_{\xi}^{2}d\xi^{2} + h_{\eta}^{2}d\eta^{2} = d\lambda^{2} + d\mu^{2}$ , with  $E = h_{\xi}^{2}$ ;  $G = h_{\eta}^{2}$ ;  $F = \frac{\partial x}{\partial \xi}\frac{\partial x}{\partial \eta} + \frac{\partial y}{\partial \xi}\frac{\partial y}{\partial \eta} + \frac{\partial z}{\partial \xi}\frac{\partial z}{\partial \eta} = 0$ - due to the orthogonality condition, where :  $\int_{\Sigma} \left( \partial x_{j} \right)^{2} \left( \frac{1}{1 - 2} \right) \left( x_{j} = x, y, z_{j} \right)$ 

$$\begin{split} \mathbf{h}_{i} &= \sqrt{\sum_{j=1}^{3} \left( \frac{\partial q_{i}}{\partial q_{i}} \right)^{2}}; \ (\mathbf{I} = \mathbf{I}, \mathbf{J}); \ \left\{ \begin{array}{l} q_{i} = \xi, \eta, \zeta_{0} \\ q_{i} = \xi, \eta, \zeta_{0} \end{array} \right\}, \text{having:} \\ \mathbf{h}_{\xi} &= \sqrt{\sum_{j=1}^{3} \left( \partial x_{j} / \partial \xi \right)^{2}} \\ \mathbf{h}_{\eta} &= \sqrt{\sum_{j=1}^{3} \left( \partial x_{j} / \partial \eta \right)^{2}} \\ \mathbf{h}_{\zeta} &= f_{2} \left( \xi, \eta, \zeta_{0i} \right) ; \\ \mathbf{h}_{\zeta} &= \sqrt{\sum_{j=1}^{3} \left( \partial x_{j} / \partial \zeta \right)^{2}} \\ \mathbf{h}_{\zeta = \zeta_{0i}} &= f_{3} \left( \xi, \eta, \zeta_{0i} \right) ; \end{split}$$

(with:  $J = \frac{D(\mathbf{x}, \mathbf{y}, \mathbf{z})}{D(\xi, \eta, \zeta)} \neq 0$ ; J – the Jacobian determinant of

this change of variables; J = 0 – gives the space curves representing the entropy singularities), so resulting on the respective "i" isentropic surface:

$$\begin{split} V_{\xi i} &= \frac{1}{h_{\xi}} \frac{\partial \Phi_{i}}{\partial \xi} = \frac{\partial \Phi_{i}}{\partial \lambda}; V_{\eta i} = \frac{1}{h_{\eta}} \frac{\partial \Phi_{i}}{\partial \eta} = \frac{\partial \Phi_{i}}{\partial \mu}; V_{\zeta i} = \frac{1}{h_{\zeta}} \frac{\partial \Phi_{i}}{\partial \zeta} = 0; \\ \Rightarrow & h_{\xi} V_{\xi i} = \frac{\partial \Phi_{i}}{\partial \xi} \quad ; \quad h_{\eta} V_{\eta i} = \frac{\partial \Phi_{i}}{\partial \eta} \quad ; \\ \frac{\partial (h_{\eta} V_{\eta i})}{\partial \xi} &= \frac{\partial}{\partial \xi} \frac{\partial \Phi_{i}}{\partial \eta} = \frac{\partial^{2} \Phi_{i}}{\partial \xi \partial \eta}; \quad \frac{\partial (h_{\xi} V_{\xi i})}{\partial \eta} = \frac{\partial}{\partial \eta} \frac{\partial \Phi_{i}}{\partial \xi} = \frac{\partial^{2} \Phi_{i}}{\partial \eta \partial \xi} \\ \text{So the relation } \Omega_{\zeta i} = 0 \text{ leads to: } \frac{\partial^{2} \Phi_{i}}{\partial \xi \partial \eta} - \frac{\partial^{2} \Phi_{i}}{\partial \eta \partial \xi} = 0, \end{split}$$

representing a true relation – Schwarz' theorem for the functions of two variables (the so-called theorem of "the equality of the mixed derivatives of the second order", they differing as to the order of differentiation). This relation proves that  $\Omega_{\zeta i} = 0$  and the existence of a 2-D "quasi-potential" function  $\Phi_i$  so that:

$$\mathbf{V}_{\xi i} = \mathbf{V}_{\lambda i} = \frac{\partial \Phi_i}{\partial \lambda} \quad ; \quad \mathbf{V}_{\eta i} = \mathbf{V}_{\mu i} = \frac{\partial \Phi_i}{\partial \mu}$$

the entropy gradient vector  $\nabla S_i$  being normal to the new introduced isentropic surfaces  $\zeta = \zeta_{0i}$ . This is evidently, taking into account Crocco's equation for an isoenergetic non-isentropic (rotational) steady gas flow:  $\mathbf{\Omega} \times \mathbf{V} = \mathbf{T} \nabla \mathbf{S}$ (see [5]), where T is the absolute local static temperature. Thus, introducing the scalar quasi-potential function  $\Phi_i(M)$  $= \Phi_i(x, y, z)$ :  $V_i = \nabla \Phi_i$ , the vector ODE of motion, joined to the continuity and the physical  $(p/\rho^{\gamma} = K_i)$  ones (and taking into consideration the local speed of sound  $a_i$ definition), enables to determine the velocity vector  $V_i$ from Steichen's equation (see [6]), usually a PDE (improperly called now the "velocity potential equation", taking into account that there is a vector  $\Omega_i \neq 0$ ; the flow being rotational, more appropriate would be the term "velocity quasi-potential equation"). In the physical equation K<sub>i</sub> is the constant of the respective "i" isentropic surface.

# 4 The velocity "quasi-potential" equation for any steady flow of a compressible fluid

This vector equation may be written in a symbolic form:

$$\nabla \mathbf{V}_{i} = \frac{\mathbf{V}_{i}}{2a_{i}^{2}} \nabla (\mathbf{V}_{i}^{2}), \text{ with } \mathbf{V}_{i} = \nabla \Phi_{i}, \text{ but } \mathbf{\Omega}_{i} \neq 0; \text{ } \mathbf{V}_{i} = |\mathbf{V}_{i}|;$$
  
$$\Rightarrow \Delta \Phi_{i} = \frac{\nabla \Phi_{i}}{2a_{i}^{2}} \nabla [(\nabla \Phi_{i})^{2}] \text{ or } \Delta \Phi_{i} = \frac{1}{a_{i}^{2}} (\nabla \Phi_{i} \cdot \nabla)^{(2)} \Phi_{i},$$

where, in a Cartesian coordinate system Oxyz:

(

$$\Delta \Phi_{i} = \frac{\partial^{2} \Phi_{i}}{\partial x^{2}} + \frac{\partial^{2} \Phi_{i}}{\partial y^{2}} + \frac{\partial^{2} \Phi_{i}}{\partial z^{2}} , \text{ and, symbolically:}$$

$$(\nabla \Phi_{i} \cdot \nabla)^{(2)} \Phi_{i} = \left(\frac{\partial \Phi_{i}}{\partial x} \cdot \frac{\partial}{\partial x} + \frac{\partial \Phi_{i}}{\partial y} \cdot \frac{\partial}{\partial y} + \frac{\partial \Phi_{i}}{\partial z} \cdot \frac{\partial}{\partial z}\right) \Phi_{i}$$

and expanding the symbolic expression in the brackets:  $\int_{-\infty}^{\infty} e^{-\frac{1}{2}x^2} dx$ 

$$(\nabla \Phi_{i} \cdot \nabla)^{(2)} \Phi_{i} = \left[ \left( \frac{\partial \Phi_{i}}{\partial x} \right)^{2} \frac{\partial^{2}}{\partial x^{2}} + \left( \frac{\partial \Phi_{i}}{\partial y} \right)^{2} \frac{\partial^{2}}{\partial y^{2}} + \left( \frac{\partial \Phi_{i}}{\partial z} \right)^{2} \frac{\partial^{2}}{\partial z^{2}} \right]$$
$$+ 2 \frac{\partial \Phi_{i}}{\partial x} \frac{\partial \Phi_{i}}{\partial y} \frac{\partial^{2}}{\partial x \partial y} + 2 \frac{\partial \Phi_{i}}{\partial y} \frac{\partial \Phi_{i}}{\partial z} \frac{\partial^{2}}{\partial y \partial z} + 2 \frac{\partial \Phi_{i}}{\partial z} \frac{\partial \Phi_{i}}{\partial z} \frac{\partial^{2}}{\partial z \partial x} \right] \Phi_{i}$$

the speed of sound  $a_i$  being given by the energy equation:

$$a_{i}^{2} = \left(\frac{\mathrm{d}p}{\mathrm{d}\rho}\right)_{\mathrm{S=S_{0i}}} = \frac{\gamma - 1}{2} \left[ \mathrm{W}^{2} - \left(\frac{\partial \Phi_{i}}{\partial x}\right)^{2} - \left(\frac{\partial \Phi_{i}}{\partial y}\right)^{2} - \left(\frac{\partial \Phi_{i}}{\partial z}\right)^{2} \right],$$

all the points at which is satisfied the previous PDE of the velocity potential  $\Phi_i$  belonging to a certain "i" isentropic surface. This new equation is identical to the velocity potential equation (see [6]), written for a potential flow only. In a triorthogonal smart intrinsic coordinate system  $O\xi\eta\zeta$  tied to these surfaces (or  $O\lambda\mu\nu$ , with  $\lambda$ ,  $\mu$ ,  $\nu$  – lengths of orthogonal arcs, with  $\lambda$  and  $\mu$  contained in the local tangent plane and  $\nu$  directed upon the normal) Laplace's operator  $\Delta$  is given by the general expression below (the function  $\Phi_i$  depending on  $\xi$ ,  $\eta$  and  $\zeta_{0i}$ , or on  $\lambda$ ,  $\mu$  and  $\nu_{0i}$ ):

$$\begin{split} \Delta \Phi_{i} &= \frac{1}{h_{\xi}h_{\eta}h_{\zeta}} \Bigg| \frac{\partial}{\partial\xi} \Bigg( \frac{h_{\eta}h_{\zeta}}{h_{\xi}} \frac{\partial \Phi_{i}}{\partial\xi} \Bigg) + \frac{\partial}{\partial\eta} \Bigg( \frac{h_{\zeta}h_{\xi}}{h_{\eta}} \frac{\partial \Phi_{i}}{\partial\eta} \Bigg) \\ &+ \frac{\partial}{\partial\zeta} \Bigg( \frac{h_{\xi}h_{\eta}}{h_{\zeta}} \frac{\partial \Phi_{i}}{\partial\zeta} \Bigg) \Bigg] \quad ; \quad \text{but} \quad V_{\zeta i} = 0 \quad \Rightarrow \quad \frac{\partial \Phi_{i}}{\partial\zeta} = 0 \quad ; \\ \Delta \Phi_{i} &= \frac{1}{h_{\xi}h_{\eta}h_{\zeta}} \Bigg[ \frac{\partial}{\partial\xi} \Bigg( \frac{h_{\eta}h_{\zeta}}{h_{\xi}} \frac{\partial \Phi_{i}}{\partial\xi} \Bigg) + \frac{\partial}{\partial\eta} \Bigg( \frac{h_{\zeta}h_{\xi}}{h_{\eta}} \frac{\partial \Phi_{i}}{\partial\eta} \Bigg) \Bigg] \Bigg|_{\zeta = \zeta_{0i}}, \text{ or } \\ \Delta \Phi_{i} &= \frac{1}{h_{\zeta}} \Bigg[ \frac{1}{h_{\eta}} \frac{\partial}{\partial\lambda} \Bigg( h_{\eta}h_{\zeta} \frac{\partial \Phi_{i}}{\partial\lambda} \Bigg) + \frac{1}{h_{\xi}} \frac{\partial}{\partial\mu} \Bigg( h_{\zeta}h_{\xi} \frac{\partial \Phi_{i}}{\partial\mu} \Bigg) \Bigg] \Bigg|_{\nu = v_{0i}}, \end{split}$$

e.g.: the  $\lambda$  arcs taken along the streamlines and the  $\mu$  arcs – along the equi(iso)-"quasi-potential" lines  $\equiv$  intersection lines of the isentropic surfaces with the iso-quasi-potential ones, these being normal to the local velocity **V**. Similarly, in Steichen's equation – a nonlinear PDE of the 2nd order in three variables –  $\xi$ ,  $\eta$ ,  $\zeta$  (now written for any rotational flow –  $\Omega \neq 0$ , but *on the* "i" *isentropic surfaces*  $\zeta = \zeta_{0i}$ ) disappear all the terms containing the partial derivative about  $\zeta$ of the potential function  $\Phi_{i}$ ,  $(\partial \Phi_i/\partial \zeta)$ , and also its derivatives with respect to  $\xi$ ,  $\eta$  and  $\zeta$ , thus being obtained a nonlinear PDE of the second order in only two variables –  $\xi$ ,  $\eta$  – the "velocity quasi-potential equation" (see [1] – [3]), which was thoroughly treated in a recent paper ([7]).

5 The interdependence between the 2-D "quasi-potential" function and the 2-D "quasi-stream" scalar function (Selescu); the orthogonality of their gradient lines Taking into account the previously established relations (also see subsection 1.2 in [1], [3] and section 2 in [2]):

$$\begin{split} \mathbf{V}_{\xi i} &= \frac{1}{\mathbf{h}_{\xi}} \frac{\partial \Phi_{i}}{\partial \xi} = \frac{\partial \Phi_{i}}{\partial \lambda} \quad ; \quad \mathbf{V}_{\eta i} = \frac{1}{\mathbf{h}_{\eta}} \frac{\partial \Phi_{i}}{\partial \eta} = \frac{\partial \Phi_{i}}{\partial \mu} \quad ; \\ \mathbf{V}_{\zeta i} &= \frac{1}{\mathbf{h}_{\zeta}} \frac{\partial \Phi_{i}}{\partial \zeta} = 0 \quad \text{and that} \quad \mathbf{V}_{i} = \nabla \Phi_{i} = \frac{1}{\rho} \nabla \times \boldsymbol{\Psi}_{ci} \quad , \end{split}$$

one gets the searched for interdependence  $\Phi_i \leftrightarrow \Psi_{ci}$ :

$$\begin{split} \mathbf{V}_{\xi i} &= \frac{1}{h_{\xi}} \frac{\partial \Phi_{i}}{\partial \xi} = \frac{1}{\rho h_{\eta}} \frac{\partial \Psi_{d}}{\partial \eta}; \\ \mathbf{V}_{\eta i} &= \frac{1}{h_{\eta}} \frac{\partial \Phi_{i}}{\partial \eta} = -\frac{1}{\rho h_{\xi}} \frac{\partial \Psi_{d}}{\partial \xi}, \\ \text{or:} \quad \mathbf{V}_{\xi i} &= \frac{\partial \Phi_{i}}{\partial \lambda} = \frac{1}{\rho} \frac{\partial \Psi_{d}}{\partial \mu} \quad ; \quad \mathbf{V}_{\eta i} = \frac{\partial \Phi_{i}}{\partial \mu} = -\frac{1}{\rho} \frac{\partial \Psi_{d}}{\partial \lambda} \quad . \end{split}$$

Let us check the orthogonality of the gradient lines above:  $\nabla \Phi_i \cdot \nabla \Psi_{ci} = (\partial \Phi_i / \partial \lambda)(\partial \Psi_{ci} / \partial \lambda) + (\partial \Phi_i / \partial \mu)(\partial \Psi_{ci} / \partial \mu) = 0$ ; therefore these vector lines are orthogonal each other.

6 A simple example: the 3-D conical flow

Let us consider further as a simple application the case of a general 3-D conical supersonic flow, with its two well-known particular cases of 2-D conical flow: axisymmetric (Busemann–Taylor–Maccoll) and plane (Prandtl –Meyer). The isentropic surfaces are the conical ones with the vertex in the cone tip and the streamlines as directrices, as a result of the fact that the entropy is constant along both the streamline and any half-straight line starting from the tip. Let be: (x, y, z) – the Cartesian coordinates;  $(R, \theta, \omega)$  – the classical spherical coordinates;  $(R, \varphi, \chi)$  – the generalized spherical (conical) coordinates – *intrinsic*;  $x = Rf_1(\theta, \omega) = R\cos\theta$ ;  $y = Rf_2(\theta, \omega) = R\sin\theta\cos\omega$ ;  $z = Rf_3(\theta, \omega) = R\sin\theta\sin\omega$ , and analogously (since x, y, z may be expressed as products of separated functions for any kind of spherical coordinates):  $x = Rf_{1G}(\varphi, \chi) = Rf_1(\varphi)g_1(\chi)$ ;  $y = Rf_{2G}(\varphi, \chi) = Rf_2(\varphi)g_2(\chi)$ ;  $z = Rf_{3G}(\varphi, \chi) = Rf_3(\varphi)g_3(\chi)$ ; (also see [8], [9]), satisfying the following condition on the Jacobian J (see subsection 1.2 in [1], [3] and section 2 in [2]):

$$J = \frac{D(\mathbf{x}, \mathbf{y}, \mathbf{z})}{D(\mathbf{R}, \varphi, \chi)} = \mathbf{R}^{2} \begin{vmatrix} \mathbf{f}_{1}g_{1} & \mathbf{f}_{2}g_{2} & \mathbf{f}_{3}g_{3} \\ \mathbf{f}_{1\varphi}'g_{1} & \mathbf{f}_{2\varphi}'g_{2} & \mathbf{f}_{3\varphi}'g_{3} \\ \mathbf{f}_{1}g_{1\chi}' & \mathbf{f}_{2}g_{2\chi}' & \mathbf{f}_{3}g_{3\chi}' \end{vmatrix} = \\ \mathbf{R}^{2}\mathbf{f}_{1}g_{1}f_{2}g_{2}f_{3}g_{3} \begin{vmatrix} \mathbf{1} & \mathbf{1} & \mathbf{1} \\ (\ln|\mathbf{f}_{1}|)_{\varphi}' & (\ln|\mathbf{f}_{2}|)_{\varphi}' & (\ln|\mathbf{f}_{3}|)_{\varphi}' \\ (\ln|\mathbf{g}_{1}|)_{\chi}' & (\ln|\mathbf{g}_{2}|)_{\chi}' & (\ln|\mathbf{g}_{3}|)_{\chi}' \end{vmatrix} \neq 0;$$

J=0- gives entropy's S singularities (see subsection 1.3 in [1], [3] and section 5 in [2]) for a cone at incidence: a pair of half-straight lines – two nodal logarithmic singularities, and a half-straight line – a false saddle point type singularity; sphere(S):  $x^2 + y^2 + z^2 = C_0^2$ ; (R = C<sub>0</sub>;)  $\cos^2 \theta + \sin^2 \theta \cos^2 \omega + \sin^2 \theta \sin^2 \omega = 1$ ;

 $\begin{array}{l} (\mathbf{R} = \mathbf{C}_0;) \quad \cos^2 \theta + \sin^2 \theta \cos^2 \omega + \sin^2 \theta \sin^2 \omega = 1; \\ \text{similarly,using the generalized spherical coordinates :} \\ \mathbf{R} = \mathbf{C}_0; \quad \mathbf{f}_1^2(\varphi) \mathbf{g}_1^2(\chi) + \mathbf{f}_2^2(\varphi) \mathbf{g}_2^2(\chi) + \mathbf{f}_3^2(\varphi) \mathbf{g}_3^2(\chi) = 1; \\ \text{cone } (\mathbf{C}_1): \quad \mathbf{F}_1(\mathbf{y}/\mathbf{x}, \mathbf{z}/\mathbf{x}) = 0; \end{array}$ 

For  $(c_1)$ ,  $F_1[\frac{f_2(\varphi)g_2(\chi)}{f_1(\varphi)g_1(\chi)}, \frac{f_3(\varphi)g_3(\chi)}{f_1(\varphi)g_1(\chi)}] = 0$ ; or  $\begin{cases} F_1[f_4(\varphi)g_4(\chi), f_5(\varphi)g_5(\chi)] = F_1(\varphi) = 0, \\ \text{with the solution : } \varphi = C_1; \end{cases}$ cone  $(C_2)$ :  $F_2(y/x, z/x) = 0;$   $F_2[\frac{f_2(\varphi)g_2(\chi)}{f_1(\varphi)g_1(\chi)}, \frac{f_3(\varphi)g_3(\chi)}{f_1(\varphi)g_1(\chi)}] = 0; \text{ or}$  $\begin{cases} F_2[f_4(\varphi)g_4(\chi), f_5(\varphi)g_5(\chi)] = F_2(\chi) = 0, \\ \text{with the solution : } \chi = 1/C_2; \end{cases}$ 

 $h_R$ ,  $h_{\varphi}$ ,  $h_{\chi}$  are Lamé's coefficients:

$$\begin{split} \mathbf{h}_{i} &= \sqrt{\sum_{j=1}^{3} \left(\frac{\partial \mathbf{x}_{j}}{\partial \mathbf{q}_{i}}\right)^{2}} \quad ; \quad (i = \overline{1, 3}) \quad ; \quad \begin{cases} \mathbf{x}_{j} = \mathbf{x}, \mathbf{y}, \mathbf{z} \\ \mathbf{q}_{i} = \mathbf{R}, \mathbf{y}, \mathbf{z} \\ \mathbf{q}_{i} = \mathbf{q}, \mathbf{q}, \mathbf{q} \\ \mathbf{q}_{i} = \mathbf{q}, \mathbf{q} \\ \mathbf{q} \\ \mathbf{q}_{i} = \mathbf{q}, \mathbf{q} \\ \mathbf{$$

We will start further from the continuity equation (1)

for a certain 3-D steady motion, in a triorthogonal curvilinear  $q_i$  coordinate system, which becomes in the generalized spherical *intrinsic* coordinates (R,  $\varphi$ ,  $\chi$ ):  $\partial(\rho h + V) = \partial(\rho h + V) = \partial(\rho h + V)$ 

$$\frac{\partial (\rho \Pi_{\varphi} \Pi_{\chi} V_{R})}{\partial R} + \frac{\partial (\rho \Pi_{R} \Pi_{\chi} V_{\varphi})}{\partial \varphi} + \frac{\partial (\rho \Pi_{R} \Pi_{\varphi} V_{\chi})}{\partial \chi} = 0 ,$$
  
with:  $h_{R} = 1$ ;  $h_{\varphi} = R$ ;  $h_{\chi} = R f(\varphi) ,$   
 $\frac{\partial}{\partial R} \left[ \rho R^{2} f(\varphi) V_{R} \right] + \frac{\partial}{\partial \varphi} \left[ \rho R f(\varphi) V_{\varphi} \right] + \frac{\partial}{\partial \chi} \left( \rho R V_{\chi} \right) = 0 ;$   
 $\Rightarrow f(\varphi) \frac{\partial \left( \rho R^{2} V_{R} \right)}{\partial R} + R \left\{ \frac{\partial \left[ \rho f(\varphi) V_{\varphi} \right]}{\partial \varphi} + \frac{\partial \left( \rho V_{\chi} \right)}{\partial \chi} \right\} = 0 ,$ 

equation that we will divide by  $R \cdot f(\phi)$  ( $\neq 0$ ), and replacing the velocity V components along the system axes:

$$V_{R} = V = R; \quad V_{\varphi} = R\dot{\varphi}; \quad V_{\chi} = R f(\varphi) \dot{\chi}, \text{ one gets}:$$
$$\frac{1}{R} \frac{\partial \left(\rho R^{2} \dot{R}\right)}{\partial R} + \frac{1}{f(\varphi)} \left\{ \frac{\partial \left[\rho R f(\varphi) \dot{\varphi}\right]}{\partial \varphi} + \frac{\partial \left[\rho R f(\varphi) \dot{\chi}\right]}{\partial \chi} \right\} = 0.$$

Taking into account that we are studying a conical flow, and R,  $\varphi$  and  $\chi$  are independent variables, chosen so that along any streamline:  $\chi = \chi_0$ ;  $\dot{\chi} = 0$ ;  $\Rightarrow V_{\chi} = 0$ , we have:

$$\frac{1}{R}\frac{\partial[\rho R^2 R]}{\partial R} + \frac{1}{f(\varphi)}\frac{\partial[\rho Rf(\varphi)\phi]}{\partial\varphi} = 0, \text{ with } : \frac{\partial\rho}{\partial R} = \frac{\partial R}{\partial R} = 0$$
(conical flow);  $\frac{\partial R}{\partial\varphi} = 0$  (R,  $\varphi$  – independent variables);  
 $\dot{\rho} = \frac{d\rho}{dt} = \frac{\partial\rho}{\partial\varphi}\dot{\phi} + \frac{\partial\rho}{\partial\chi}\dot{\chi} = \frac{\partial\rho}{\partial\varphi}\dot{\phi} \quad ; \implies \frac{\partial\rho}{\partial\varphi} = \frac{\dot{\rho}}{\dot{\phi}} \quad ;$ 

$$\frac{\partial f(\varphi)}{\partial q} = \frac{df(\varphi)}{\partial q} = \frac{df(\varphi)}{\partial q} \frac{dt}{dt} = \frac{1}{2}\frac{df(\varphi)}{dt} = \frac{[f(\varphi)]}{2}$$

Grouping conveniently the terms and dividing by  $\rho R \neq 0$ , we obtain the form of the continuity equation, specific to the conical flows (a first order PDE in one variable  $-\varphi$ ):

dt d $\phi$   $\dot{\phi}$  dt

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$$\frac{\dot{\rho}}{\rho} + 2\frac{\dot{R}}{R} + \frac{\left[f(\varphi)\right]}{f(\varphi)} + \frac{\partial\dot{\varphi}}{\partial\varphi} = 0 \quad , \tag{5}$$

similar to Eq. (3) previously established for a certain 3-D steady flow, where:  $\xi = R$ ;  $\eta = \varphi$  and  $\zeta = \chi$ . We will try to express in other form the partial derivative  $\partial \dot{\varphi} / \partial \varphi$ :

$$\ddot{\varphi} = \frac{\mathrm{d}\dot{\varphi}}{\mathrm{d}t} = \frac{\partial\dot{\varphi}}{\partial \mathrm{R}}\dot{\mathrm{R}} + \frac{\partial\dot{\varphi}}{\partial\varphi}\dot{\varphi} + \frac{\partial\dot{\varphi}}{\partial\chi}\dot{\chi} \quad , \quad \left(\frac{\partial\dot{\varphi}}{\partial\mathrm{R}} \neq 0\right),$$

 $\dot{\phi}$  varying with respect to R, because  $\partial(R\dot{\phi})/\partial R = 0$ ,

whence it results in:  $\frac{\partial \dot{\phi}}{\partial \varphi} = \frac{\ddot{\phi}}{\dot{\phi}} - \frac{\partial \dot{\phi}}{\partial R}\frac{\dot{R}}{\dot{\phi}} - \frac{\partial \dot{\phi}}{\partial \chi}\frac{\dot{\chi}}{\dot{\phi}}$ 

Taking into account that we study a conical flow, in a special chosen coordinate system, there are the relations:

$$\frac{\partial \mathbf{V}_{\varphi}}{\partial \mathbf{R}} = \frac{\partial (\mathbf{R}\,\dot{\varphi})}{\partial \mathbf{R}} = 0 \text{ and } \dot{\chi} = 0$$

 $\partial \varphi$ 

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Developing the first one, we get:  $\dot{\phi} + R \frac{\partial \dot{\phi}}{\partial R} = 0$ .

Therefore, in the case of a certain conical flow, we have:  $\partial \dot{\phi} / \partial \mathbf{R} = - \dot{\phi} / \mathbf{R}$ . Replacing this in the expression of  $\partial \dot{\varphi} / \partial \varphi$ , one obtains:

$$\frac{\partial \dot{\varphi}}{\partial \varphi} = \frac{\ddot{\varphi}}{\dot{\varphi}} + \frac{\dot{\varphi}}{R} \frac{R}{\dot{\varphi}} = \frac{R}{R} + \frac{\ddot{\varphi}}{\dot{\varphi}} ,$$

the previous continuity Eq. (5) for conical flows getting the first integrable general form below (a second order ODE):

$$\frac{\dot{\rho}}{\rho} + 3\frac{\dot{R}}{R} + \frac{[f(\varphi)]}{f(\varphi)} + \frac{\ddot{\varphi}}{\dot{\varphi}} = 0, \text{ or } \frac{d}{dt} \left[ \ln \left| \rho R^{3} f(\varphi) \dot{\varphi} \right| \right] = 0, \quad (6)$$

valid for any conical flow (either potential or rotational), representing the new form of its continuity equation, admitting a remarkable first integral (the flow rate equation):  $P^{2}(x) = P^{2}(x) V$ 

$$\rho R^{2} f(\varphi) \varphi = Q_{i}, \text{ or } \rho R^{i} f(\varphi) V_{\varphi} = Q_{i}, \text{ or more}:$$

$$\rho R^{2} f(\varphi) V' = Q_{i} \quad (\text{with } ' = d/d\varphi) \quad (7)$$

(see example 3 at the end of subsection 1.2 in [1], [3]), the constants Q<sub>i</sub> being kept over the field surfaces  $\psi(\mathbf{R}, \varphi) = \mathbf{Q}_i$ . The quantity  $R^2 f(\phi)$  is proportional to the infinitesimal side area of the current cone ( $C_1$ ):  $\varphi = \varphi_{0i} = C_1$  (fig. 1.b in [1] – [3], [8]), between two closely near (isentropic) cones  $(C_2)$ :  $\chi = \chi_{0i} = c(S_{0i} - c_0) = 1/C_2$  (fig. 1.a in [1] – [3], [8]); the tip half-angle  $\varphi$  of this quasi-circular cone (compression simple wave) is an intrinsic coordinate. One can apply a qualitative analogy of conical isentropic sheets traces to the streamlines in an incompressible 2-D plane potential flow. In fig. 1 the flow due to a two source system is represented. A more realistic flow pattern can be obtained adding a third source. In the particular case of the axisymmetric conical flow (Busemann-Taylor-Maccoll: see [10], [11]), we have:  $\varphi = \theta$ ;  $\dot{\varphi} = \dot{\theta}$ ;  $\ddot{\varphi} = \ddot{\theta}$ ;  $f(\varphi) = f(\theta) = \sin\theta$ ;  $[f(\varphi)] = [f(\theta)] = (\sin\theta) = \cos\theta \cdot \dot{\theta}$ ,

so being obtained for the ODE (6) the first integrable form:

$$\frac{\dot{\rho}}{\rho} + 3\frac{\dot{R}}{R} + \frac{(\sin\theta)}{\sin\theta} + \frac{\ddot{\theta}}{\dot{\theta}} = 0, \text{ or : } \frac{\dot{\rho}}{\rho} + 3\frac{\dot{R}}{R} + \cot\theta \cdot \dot{\theta} + \frac{\ddot{\theta}}{\dot{\theta}} = 0$$

and respectively, for the first integral (7) (also see [8], [12]):  $\rho R^3 \sin\theta \dot{\theta} = \rho R^2 \sin\theta V_{\theta} = \rho R^2 \sin\theta \cdot V' = Q_3 = m/\pi$ . The quantity  $\pi R^2 \sin\theta$  represents just the side area of the current circular cone of vertex half-angle  $\theta$ , in the constant  $Q_3$  expression intervening the mass flow rate *m*. In case of the plane 2-D conical flow, it can be easily proved that:

$$\begin{split} \varphi &= \omega \quad ; \quad \dot{\varphi} = \dot{\omega} \quad ; \quad \ddot{\varphi} = \ddot{\omega} \quad ; \quad f(\varphi) = f(\omega) = \frac{C}{R(\omega)} \quad ; \\ \dot{\left[f(\omega)\right]} &= -\frac{C\dot{R}(\omega)}{R^{2}(\omega)} \quad ; \quad \frac{\left[f(\omega)\right]}{f(\omega)} = -\frac{\dot{R}}{R} \quad , \end{split}$$

representing the plane conical flow (with Prandtl–Meyer expansion or isentropic compression; see [13] – [17]):

$$\begin{split} \varphi &= \omega; \ ' = \frac{d}{d\varphi} = \frac{d}{d\omega}; \ f(\varphi) = f(\omega) = \cos^{\frac{\gamma+1}{\gamma-1}} \left[ \sqrt{\frac{\gamma-1}{\gamma+1}} (\omega + \omega_0) \right] \\ \left[ \ln \left| f(\varphi) \right| \right]' &= -\sqrt{\frac{\gamma+1}{\gamma-1}} \tan \left[ \sqrt{\frac{\gamma-1}{\gamma+1}} (\omega + \omega_0) \right] ; \end{split}$$

the angle of attack  $\alpha \neq 0$ ; r is the cone local radius;



- Cartesian coordinates, x - abscissa of cross section current plane Fig. 1. Simple qualitative example of "generalized spherical" smart intrinsic coordinate surfaces for the case of a circular cone at a small angle of attack (cross section) - the incompressible approximation (slender body; no shock wave), giving the relations between the new spherical coordinates (R,  $\varphi$ ,  $\chi$ ) and the Cartesian ones; the flow is due to two semi-infinite line sources along cone's axis (a) and back (b), replacing it; (a), (b), (n) – nodal and saddle lines; 1.a. the conical isentropic sheets  $\chi = \chi_{0i} = c(S_{0i} - c_0) = 1/C_2$  (a smart intrinsic coordinate tied to  $S_{0i}$  – the local specific entropy value), having as remarkable directrices: the oz axis and a circle (the solid cone trace) centered on it (both for  $C_2 = 0$ ), and a right strophoid ( $\chi = 0$ ) centered on oz axis too (for  $1/C_2 = 0$ ); c, S<sub>0i</sub>, c<sub>0</sub>>0; 1.b. the conical sheets  $\varphi = \varphi_{0j} = C_1$  (smart intrinsic coordinate), orthogonal to the conical isentropic ones:  $\nabla \overline{Y}(\overline{y}, \overline{z}) \cdot \nabla \overline{Z}(\overline{y}, \overline{z}) = 0$ , having as remarkable directrices: a Pascal's limaçon and the circle at infinity (both for  $C_1 = 0$  and centered on the oz axis).

(the isentropic surfaces are the planes parallel to that represented by the degenerate cone  $\theta = \pi/2$ , that means:  $\chi = x = x_0$ ; the  $\varphi = \varphi_0 = C_1$  intrinsic coordinate surfaces are the planes  $\varphi = \omega = \omega_0$ ), being thus proved that:

$$f(\omega) = \frac{R_0}{R(\omega)}$$
 and  $\frac{f'(\omega)}{f(\omega)} = -\frac{R'(\omega)}{R(\omega)}$ , or  $\frac{\dot{f}}{f} = -\frac{\dot{R}}{R}$ ,

so being obtained the first integrable form below:

 $\frac{\dot{\rho}}{\rho} + 2\frac{\dot{R}}{R} + \frac{\ddot{\omega}}{\dot{\omega}} = 0$ , the first integral (7) becoming:

 $\rho R^2 \dot{\omega} = \rho R V_{\omega} = \rho R V' = Q_2 = m/l$  (also see [8], [12]), where *l* is the thickness (depth) of the plane flow layer.

# 7 The flow rate equation; special and general cases of first integrability for the continuity equation in a thick stream tube The first integrability of the general (even unsteady) continuity equation (like Eqs. (2), (3)) will be treated in section 12.

The *fluid flow* across a stationary surface S is the mass m of the fluid passing across this surface in unit time (the *mass flow rate*). Thus:

$$m = \int_{S} \rho V_{n} dS = \int_{S} \mathbf{j} \mathbf{n} dS$$
,

where:  $\mathbf{n}$  – versor (unit vector) of the outward normal to the element of surface d*S*;  $\mathbf{V}_n = \mathbf{V}\mathbf{n}$  – projection of the fluid velocity on the vector  $\mathbf{n}$ ;  $\mathbf{j} = \rho \mathbf{V}$  – vector of the density of fluid flow (the mass flux density);  $\mathbf{j}\mathbf{n} = \rho \mathbf{V}\mathbf{n}$  – the normal mass flux density. It follows from the continuity equation that for a steady flow of a fluid in a pipe:

 $m = \int_{S} \rho V_n dS = \text{const.}$ 

With a steady motion, the flow of fluid through a cross section of a stream filament does not depend on the location of the cross section. For two arbitrary cross sections  $dS_1$  and  $dS_2$  of an elemental filament the following condition holds:  $\rho_1 |V_{n1}| [dS_1] = \rho_2 |V_{n2}| [dS_2]$ ,

also called "the continuity equation for the steady flow in thin stream tubes". In some very special cases, if the quantities  $\rho_i$  and  $V_{ni}$  are rigorously constant on each "j" section (sections of uniform flow of a thick stream tube - also isochoric surfaces), then the ODE above admits in these sections a first integral – the (mass) flow rate equation: , or more:  $\rho_1 |\mathbf{V}_{n1}| [S_{u1}] = \rho_2 |\mathbf{V}_{n2}| [S_{u2}] = m = \text{const.}$  $\rho_{j}|\mathbf{V}_{j}\cdot\mathbf{n}_{j}|[S_{uj}] = \rho_{j}|\mathbf{V}_{nj}|[S_{uj}] = m = \text{const.}$ Here  $[S_{uj}]$  is the stream tube cross section area in the "j" section of uniform flow  $S_{uj}$ , crossing the "i" surfaces ( $\zeta$ =  $\zeta_{0i}$ ), and  $V_{nj}$  is the component of  $V_j$  taken on the local normal to the respective "j" section of uniform flow. This is a special first integral (obtained for particular cases) of the continuity equation for a steady flow of a compressible fluid. Simple examples of such flows are: - the flow due to a source of constant mass flow rate (plane and spherical – in a conical tubular domain); - the plane 2-D conical supersonic flow (Prandtl-Meyer

flow): with expansion, or with isentropic compression in a streamlined channel (a reverse of the expansion flow); - the axisymmetric 2-D conical supersonic flow (Busemann -Taylor-Maccoll flow) with isentropic compression:

1. downstream of an axisymmetric attached conical shock wave – around an infinite circular solid cone without incidence (between the shock wave and the solid cone), and

2. upstream of an axisymmetric conical shock wave with the same tip as the conical simple compression waves fan – inside Busemann's nozzle connecting two parallel and uniform co-axial supersonic flows.

The first integral (8) is valid over the volume inside any stream tube surface (a tubular surface composed of streamlines as generatrices and a certain closed curve contained in a "j" section of uniform flow as directrix). The most general case for which the continuity equation in its global form:  $\oint \mathbf{oVnd} \mathcal{L} = \oint \mathbf{ind} \mathcal{L} = \int \mathbf{ind} S + \int \mathbf{ind} S = 0$ 

$$\oint_{\mathcal{F}} \rho \mathbf{V} \mathbf{n} d\mathcal{F} = \oint_{\mathcal{F}} \mathbf{j} \mathbf{n} d\mathcal{F} = \int_{S_{in}} \mathbf{j} \mathbf{n} dS + \int_{S_{out}} \mathbf{j} \mathbf{n} dS = 0 \quad ,$$

admits as first integral a similar one to Eq. (8) (the same for the compressible flow in thin as well thick stream tubes) is:

$$|\mathbf{j}\mathbf{n}| = \rho |\mathbf{V} \cdot \mathbf{n}| = \rho |\mathbf{V}_n| = \text{const.}(\mathbf{j})$$

on some sections  $S_j$  crossing a closed surface  $\mathscr{S}$  of a thick stream tube (the lids of this thick stream tube, delimiting a control volume). That introduces the surfaces  $S_j$  of *iso*-normal mass flux *density*  $(\mathbf{jn})_j$  (a scalar quantity):  $(|\mathbf{jn}|_j = (\rho | \mathbf{V} \cdot \mathbf{n}|)_j = (\rho | \mathbf{V}_n |)_j = m/[S_j] = \text{const.}(j)$ , (9) whose envelopes are just the "j" sections of uniform flow – similarly to Huygens' principle (see [18]) explaining the propagation of a wavefront due to the spherical wavelets emanating from every point on the wavefront at the prior instant.

## 8 A method for determining the integral surfaces (lines) of the continuity equation in a thick stream tube for the 2-D flows

Assuming the local values of density  $\rho$  and velocity V as being given (known) all over the respective domain of plane flow, the problem solution is given by a system of two ODEs, expressing the constancy of **jn** over (along) the integral "j" surfaces (lines) – see Eq. (9), and the length of an elementary arc ds (s being the intrinsic curvilinear coordinate) in the Cartesian coordinates (x, y), respectively, where  $\rho(x, y), V_x(x, y)$  and  $V_y(x, y)$  are given functions of x and y:  $\int \rho V_x (dy/ds) - \rho V_y (dx/ds) = \text{const.}(j) = k_j$ ;  $\int (dx/ds)^2 + (dy/ds)^2 = 1$ ,

the normal form (the solution) of which is given by:

$$\begin{cases} \frac{dx}{ds} = \frac{1}{\rho(V_x^2 + V_y^2)} \left[ -k_j V_y \pm V_x \sqrt{\rho^2 (V_x^2 + V_y^2)^2 - k_j^2} \right]; \\ \frac{dy}{ds} = \frac{1}{\rho(V_x^2 + V_y^2)} \left[ k_j V_x \pm V_y \sqrt{\rho^2 (V_x^2 + V_y^2)^2 - k_j^2} \right]. \end{cases}$$

For the case of an axisymmetric flow, replacing y by r in a certain meridian plane in cylindrical coordinates (x, r), the above solution remains valid, so obtaining the meridian line of the axisymmetric integral "j" surfaces. The sign (±) before the square root in the solution (dx/ds, dy/ds) means that through any flow's point are passing two "j" surfaces (lines) of *iso*-normal mass flux *density*. The quantity under this square root must be positive or zero, the physical significance of the root being the tangential mass flux density. If this quantity is just zero, a single solution remains, representing the "j" surface (line) of normal uniform flow passing through the same flow's point. Along this line we have  $|V_i| = k_i/\rho_i$ , the solution not depending on  $\rho$ :

$$\begin{cases} \left(\frac{dx}{ds}\right)_{j} = -\frac{k_{j}V_{jy}}{\rho_{j}\left(V_{jx}^{2} + V_{jy}^{2}\right)} = -\frac{k_{j}V_{jy}}{\rho_{j}\left|V_{j}\right|^{2}} = -\frac{V_{jy}}{\left|V_{j}\right|} \quad ; \\ \left(\frac{dy}{ds}\right)_{j} = \frac{k_{j}V_{jx}}{\rho_{j}\left(V_{jx}^{2} + V_{jy}^{2}\right)} = \frac{k_{j}V_{jx}}{\rho_{j}\left|V_{j}\right|^{2}} = \frac{V_{jx}}{\left|V_{j}\right|} \quad . \end{cases}$$

Therefore through any 2-D flow's point three integral "j" lines (two  $S_j$  and one normal  $S_{uj}$ ) are passing. If the root has a constant value  $k_{1j}$ , one gets a pair of solutions (±), representing other two "j" lines of uniform flow  $S_{uj}$ , this time not normal, like the half-straight lines for the plane (Prandtl–Meyer) and axisymmetric (Busemann–Taylor– Maccoll) conical flows. In these cases more conveniently is to use the polar plane (R,  $\omega$ ), and polar spherical (R,  $\theta$ ) coordinates (see section 6) instead of the Cartesian ones. Furthermore, this method can be extended for the 3-D flows, in the new smart intrinsic orthogonal coordinates ( $\xi$ ,  $\eta$ ,  $\zeta$ ), or ( $\lambda$ ,  $\mu$ ,  $\nu$ ), used in sections 1 – 5 of this work, in order to find one (that in the plane ( $\xi$ ,  $\eta$ ), or ( $\lambda$ ,  $\mu$ ) tangent to the (**V**,  $\Omega$ ) surface) of the two orthogonal lines which determine the local integral "j" surface (**jn** = const.) for the continuity equation.

# 9 A simple example: the plane flow due to a source of constant mass flow rate

Let us consider as simple example the case of a plane flow of a compressible fluid due to a source of constant mass flow rate (between two radii – half-straight lines). The flow is not a conical one, but is closely related to this (circular cylindrical), the main physical quantities (velocity, static pressure, static temperature, density, entropy) being kept constant on some concentric circles (co-axial cylinders). The surfaces  $S_i$  of *iso*-normal mass flux *density* (**jn**)<sub>i</sub> are the families of circles (circular cylinders) with the same diameter  $d_i$  and passing through the respective source trace (origin, or the Oz axis – a singularity for **jn**:  $(jn)_0 = 0$ ), given by the Cartesian equation below:  $(x - d_j/2 \cdot \cos\theta_k)^2 + (y - d_j/2 \cdot \sin\theta_k)^2 = (d_j/2)^2$ , or  $x^2 + y^2 - d_j(x \cos\theta_k + y \sin\theta_k) = 0$ , with  $\theta_k \in (0, 2\pi)$ , or by the polar equation  $R = d_i \cos(\theta - \theta_k)$ , with  $\theta, \theta_k \in (0, 2\pi)$ . Their envelopes (the surfaces  $S_{uj}$  of uniform flow) are the concentric circles (co-axial circular cylinders) of radius  $R = d_i$ , having the source trace as center (the Oz axis as symmetry axis), given by the Cartesian equation:  $x^2 + y^2 - d_i^2 = 0$ .

But this flow (either subsonic or supersonic) takes place outside the circle (circular cylinder)  $R = R_c$  only ( $R_c$  is the critical radius, on which V = a = c – the critical speed at the considered flow point), so that one must have:  $d_j \ge R_c$ ; ( $R < R_c - a$  no-motion region), by physical reasons (see fig. 2). Therefore the surfaces  $S_j$  of *iso*-normal mass flux *density* (**jn**)<sub>j</sub> are the families of circles (circular cylinders) above, except for their regions inside the circle (circular cylinder) of equation  $R = R_c$ , so including the singularity in origin for the *iso*-normal mass flux *density* (**jn**)<sub>0</sub>. The flow rate Eqs. (8), (9) take in this case the form:  $2\pi d_i$ (**jn**)<sub>i</sub> = m = const.,



Fig. 2. The pattern of the iso-normal mass flux density surfaces and the sections of uniform flow for a simple plane flow of a compressible fluid due to a source of constant mass flow rate m O – the source trace in the xOy plane (orthogonal projection of the source filament directed on the Oz-axis);  $S_i$  – the lines of *iso*-normal mass flux *density*  $|\mathbf{jn}| = \rho |\mathbf{Vn}| = \rho |\mathbf{V_n}| = \text{const}(\mathbf{j}) - \text{circles}$ passing through O (traces in the xOy plane of the right circular cylinders "j" with their axes parallel to the Oz-axis and of diameter  $d_i$ ), given by the polar equation:  $R = d_i \cos(\theta - \theta_k)$ , with  $\theta_k \in (0, 2\pi)$ ; along a certain circle  $\theta_k$  takes a constant value;  $\theta$  is the current polar angle;  $S_{ui}$  – the lines of uniform flow (on which  $\rho = \rho_j = const_1(j)$  and  $V_n = V_{nj} = const_2(j)$ ) - circles centered in O (traces in the xOy plane of the right circular cylinders "j" with their axes directed on the Oz-axis and of diameter 2d<sub>i</sub> – the envelope sheets of the cylinders "j" of iso-normal mass flux density), given by the polar eq.  $R = d_i$ . O is a singular point for |jn|. But this flow (either subsonic or supersonic) takes place outside the "critical" cylinder (circle) only:  $R \ge R_c = m/(2\pi\rho_c c)$  – the critical radius, with  $\rho_c c = \{\gamma p_0 \rho_0 [2/(\gamma + 1)]^{(\gamma + 1)/(\gamma - 1)}\}^{1/2}$  – the critical *iso*-normal mass flux *density*  $(|\mathbf{jn}|)_c$ , so that one must have:  $d_i \ge R_c$ , by physical reasons;  $R < R_c - a$  no-motion region (a solid nucleus). The lengths of radial arrowheads are proportional to density **[jn**]. Through any "j" flow's point three integral surfaces (lines) are passing: one of normal uniform flow  $(S_{uj})$  and other two of isonormal mass flux *density* ( $S_i$ ), all for the same value of  $|j\mathbf{n}|$ . These results can be generalized for a spherical source.

# 10 New global forms for the flow rate equation and continuity equation, using the 3-D "stream function" vector; analogy with Helmholtz' 2nd theorem on vortices

Considering  $\Psi_c$  definition and applying a Stokes' formula, we get for the normal flux of curl  $\Psi_c$  through the surface  $S_j$ :  $m = \int_{S_i} \mathbf{j} \mathbf{n} dS = \int_{S_i} (\nabla \times \Psi_c) \cdot \mathbf{n} dS = \oint_{I_i} \Psi_c \cdot d\mathbf{R} = (\Gamma_{\Psi_c})_{I_j} = \text{const.}$ 

- the circulation of the vector  $\Psi_c$  along the closed curve  $l_j$  bordering the surface  $S_j$ , where, this time  $S_j$  is an arbitrary cross section of the thick stream tube above – *a new global form of the flow rate equation*, leading to:  $\oint \mathbf{ind} \mathcal{L} = \int \mathbf{ind}S + \int \mathbf{ind}S + \int \mathbf{ind}S = 0$ , or:

$$\int_{S_{in}} (\nabla \times \Psi_{c}) \cdot \mathbf{n} dS + 0 \text{ (stream tube)} + \int_{S_{out}} (\nabla \times \Psi_{c}) \cdot \mathbf{n} dS$$
$$= \oint_{l_{in}} \Psi_{c} \cdot d\mathbf{R} + \oint_{l_{out}} \Psi_{c} \cdot d\mathbf{R} = (\Gamma_{\Psi_{c}})_{in} + (\Gamma_{\Psi_{c}})_{out} = 0 \quad ,$$

(d**R** is an elementary displacement vector;  $\mathscr{S}$  is defined in section 7, being composed of  $S_{\rm in}$ ,  $S_{\rm side}$  and  $S_{\rm out}$ ), so getting:  $(\Gamma_{\Psi_{\rm c}})_{\rm out} = -(\Gamma_{\Psi_{\rm c}})_{\rm in} = \text{const.}$  (10) - a new global form of the continuity equation, likeHelmholtz' 2nd theorem (see [19] – [24]) about the velocity circulation  $\Gamma_{\rm V}$  along any closed curve surrounding a vortex tube in an ideal (inviscid) fluid (see fig. 3, helping to give an alternative demonstration of the theorem above).



Fig. 3. The circulation  $\Gamma_{\Psi_c}$  of the 3-D "stream function" vector  $\Psi_c$  along any complex closed curve *C* bordering a cut surface  $S_{side}$  of a thick stream tube of a compressible perfect fluid flow,  $\Psi_c$  defined by:  $\nabla \times \Psi_c = \rho V (\Psi_c \neq \nabla G)$  – the analogy with the circulation  $\Gamma_V$  of the velocity V along any closed curve surrounding a vortex tube, according to Helmholtz' 2nd theorem about vortices in a perfect fluid, which states that  $\Gamma_V$  is constant along this vortex tube and constant in time also;  $\mathscr{S}$  – a closed *surface* of control volume of the tube (*bordered* by the lids  $S_{in}$  and  $S_{out}$ ), so that *it* includes both lids ( $S_{in}$  and  $S_{out}$ ):  $S_{in}$ ,  $S_{out} \subset \mathscr{S}$ ;  $\mathscr{S} \equiv (S_{in} \cup S_{side} \cup S_{out})$ ; AB – a cut on the tube surface, between  $l_{in}$  and  $l_{out}$  (the border curves of  $S_{in}$  and  $S_{out}$ ), so that the closed curve  $C \equiv (l_{in} \cup AB \cup l_{out} \cup BA)$  surrounds a simply-connected domain  $S_{side}$ , allowing to apply a Stokes' formula for the normal flux of  $\nabla \times \Psi_c$  through the surface  $S_{side}$ .

So:  $\int_{Sside} \mathbf{j} \mathbf{n} dS = \int_{Sside} (\nabla \times \Psi_c) \mathbf{n} dS = \int_C \Psi_c d\mathbf{R} = \int_{in} \Psi_c d\mathbf{R} + \int_{AB} \Psi_c d\mathbf{R}$ +  $\int_{out} \Psi_c d\mathbf{R} + \int_{BA} \Psi_c d\mathbf{R} = (\Gamma_{\Psi_c})_{in} + (\Gamma_{\Psi_c})_{out} = 0$ , leading to Eq. (10).

#### 11 A first integrability case for the system of two equations for the steady rotational flow of an inviscid compressible fluid

The general (unsteady, rotational) flow of a viscous Newtonian compressible fluid is described by the system

$$\begin{cases} \partial \mathbf{V} / \partial t + \nabla (\mathbf{V}^2 / 2) + \mathbf{\Omega} \times \mathbf{V} = \mathbf{f} - [\nabla p - \mu_1 \Delta \mathbf{V} \\ -(\mu_2 + \mu_1 / 3) \cdot \nabla (\nabla \mathbf{V})] / \rho, \quad \mu_1 = ct.; \quad \mu_2 = ct.; \\ \partial \rho / \partial t + \nabla (\rho \mathbf{V}) = 0 \quad , \end{cases}$$

in first equation (Navier–Stokes) using classical symbols (also see [24], p. 193 – the viscous Crocco–Vázsonyi eq.). One must search for first integrals for this system of simultaneous equations above (motion and continuity). For the case of a steady flow the system becomes:

 $\begin{cases} \nabla (\mathbf{V}^2/2) + \mathbf{\Omega} \times \mathbf{V} = \mathbf{f} - [\nabla p - \mu_1 \Delta \mathbf{V} - (\mu_2 + \mu_1/3) \cdot \nabla (\nabla \mathbf{V})] / \rho; \\ \nabla (\rho \mathbf{V}) = 0 \end{cases}$ 

Both equations admit first integrals: the first one – along the intersection lines of the zero-work surfaces (for the nonconservative force densities – see [1], [3], [25]) with the isentropic ones (see the same); the second one – over the sections of uniform flow (if they exist), and over the surfaces of *iso*-normal mass flux *density*. These integrals do not represent the system's first integrals, due to their different definition (with no intersection domain). In the case of an inviscid fluid flow the system above becomes simpler:  $\nabla(\mathbf{V}^2/2) + \mathbf{\Omega} \times \mathbf{V} = \mathbf{f} - \nabla p/\rho$ ;  $\nabla(\rho \mathbf{V}) = 0$ .

The first integrals for these equations are given: by Eq. (11) in [1], [3], valid over any  $(\mathbf{V}, \boldsymbol{\Omega})$  isentropic "i" surface (a), and by Eq. (5) in [9], valid over any polytropic surface (b) - a more general case, including the previous one, and, on the other hand, by the flow rate equation (8), valid over the "j" sections of uniform flow, on which:  $\rho = \text{const.}$  and  $|V_n| = \text{const.}$  as well (c), and over the surfaces S<sub>j</sub> of *iso*-normal mass flux *density*:  $\rho |\mathbf{V} \cdot \mathbf{n}| = \rho |\mathbf{V}_{\mathbf{n}}| = \text{const.}$ (d), whose envelope sheets are just the sections of uniform flow of a thick stream tube (if they exist), respectively. For an incompressible fluid flow the sheet families (c) and (d) coincide: (c)  $\equiv$  (d). On the intersection lines of these two pair of surface families:  $[(a) \cap (c)]$ , or  $[(a) \cap (d)]$ , and:  $[(b) \cap (c)]$ , or  $[(b) \cap (d)]$ , one obtains the validity domain of the first integrals for the system of equations for the steady rotational flow of an inviscid compressible fluid. But, since the various polytropic surfaces passing through a certain point are in infinite number (a star of sheets),

this results in satisfying the integrability conditions for both equations over the surfaces (c) and (d) only. A similar problem of a first integrability case for a system of simultaneous vector equations occurs in MHD and was treated in subsection 2.10 in [26]. Besides the two equations above, the "magnetic induction" one must be considered.

#### 12 The first approach to the new unsteady model (the complete continuity equation); two related rare cases of first integrability

Let us start from the equivalent local form of this equation  $d\rho/dt + \rho\nabla V = 0$  – the complete continuity equation (see section 1), for the unsteady flow of a compressible fluid. Dividing the equation by  $\rho \neq 0$  one obtains the form:  $(1/\rho) \cdot (d\rho/dt) + \nabla V = 0$ , or more:  $d(\ln \rho)/dt + \nabla V = 0$ . We apply further the same procedure as in section 1, using Lamé's coefficients and the total time-derivatives. In a triorthogonal system of curvilinear coordinates  $\nabla V =$ 

$$\frac{1}{\mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}}\left[\frac{\partial}{\partial\xi}\left(\mathbf{h}_{\eta}\mathbf{h}_{\zeta}\mathbf{V}_{\xi}\right)+\frac{\partial}{\partial\eta}\left(\mathbf{h}_{\xi}\mathbf{h}_{\zeta}\mathbf{V}_{\eta}\right)+\frac{\partial}{\partial\zeta}\left(\mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{V}_{\zeta}\right)\right]$$

with  $~V_{\xi}=h_{\xi}\dot{\xi}$  ;  $V_{\eta}=h_{\eta}\dot{\eta}$  ;  $V_{\zeta}=h_{\zeta}\dot{\zeta}~$  , so having  $\nabla V\!=\!$ 

$$\begin{split} & \frac{1}{\mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}} \left[ \frac{\partial}{\partial\xi} \left( \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}\dot{\xi} \right) + \frac{\partial}{\partial\eta} \left( \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}\dot{\eta} \right) + \frac{\partial}{\partial\zeta} \left( \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}\dot{\zeta} \right) \right] \\ &= \frac{1}{\mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta}} \left[ \left( \dot{\xi}\frac{\partial}{\partial\xi} + \dot{\eta}\frac{\partial}{\partial\eta} + \dot{\zeta}\frac{\partial}{\partial\zeta} \right) \left( \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta} \right) \\ &+ \mathbf{h}_{\xi}\mathbf{h}_{\eta}\mathbf{h}_{\zeta} \left( \frac{\partial\dot{\xi}}{\partial\xi} + \frac{\partial\dot{\eta}}{\partial\eta} + \frac{\partial\dot{\zeta}}{\partial\zeta} \right) \right] = \end{split}$$

and, taking into account the local time-derivative  $(\partial/\partial t)$ ,

$$\begin{aligned} \frac{1}{h_{\xi}h_{\eta}h_{\zeta}} \left[ \left( \frac{d}{dt} - \frac{\partial}{\partial t} \right) (h_{\xi}h_{\eta}h_{\zeta}) + h_{\xi}h_{\eta}h_{\zeta} \left( \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} \right) \right] \\ = \frac{1}{h_{\xi}h_{\eta}h_{\zeta}} \left( \frac{d}{dt} - \frac{\partial}{\partial t} \right) (h_{\xi}h_{\eta}h_{\zeta}) + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} \\ = \frac{d}{dt} \ln(h_{\xi}h_{\eta}h_{\zeta}) - \frac{\partial}{\partial t} \ln(h_{\xi}h_{\eta}h_{\zeta}) + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta}, \text{ or } \\ \left[ \ln(h_{\xi}\dot{h}_{\eta}h_{\zeta}) \right] - \left( \frac{\partial \ln h_{\xi}}{\partial t} + \frac{\partial \ln h_{\eta}}{\partial t} + \frac{\partial \ln h_{\zeta}}{\partial t} \right) + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} \\ = \frac{\dot{h}_{\xi}}{h_{\xi}} + \frac{\dot{h}_{\eta}}{h_{\eta}} + \frac{\dot{h}_{\zeta}}{h_{\zeta}} - \left( \frac{1}{h_{\xi}}\frac{\partial h_{\xi}}{\partial t} + \frac{1}{h_{\eta}}\frac{\partial h_{\eta}}{\partial t} + \frac{1}{h_{\zeta}}\frac{\partial h_{\zeta}}{\partial t} \right) + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} \\ \text{the complete equation becoming thus successively:} \\ \frac{d}{dt} \ln \rho + \frac{d}{dt} \ln(h_{\xi}h_{\eta}h_{\zeta}) - \frac{\partial}{\partial t} \ln(h_{\xi}h_{\eta}h_{\zeta}) + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} \\ = 0; \frac{d}{dt} \ln(\rho h_{\xi}h_{\eta}h_{\zeta}) - \frac{\partial}{\partial t} \ln(h_{\xi}h_{\eta}h_{\zeta}) + \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} = 0, \\ \text{differing from Eq. (2) by the 2nd term. In an expanded form } \\ \frac{\dot{\rho}}{\rho} + \frac{\dot{h}_{\xi}}{h_{\xi}} + \frac{\dot{h}_{\eta}}{h_{\eta}} - \left( \frac{1}{h_{\xi}}\frac{\partial h_{\xi}}{\partial t} + \frac{1}{h_{\eta}}\frac{\partial h_{\eta}}{\partial t} + \frac{1}{h_{\zeta}}\frac{\partial h_{\zeta}}{\partial t} - \left( \frac{\partial \dot{\xi}}{\partial \xi} + \frac{\partial \dot{\eta}}{\partial \eta} + \frac{\partial \dot{\zeta}}{\partial \zeta} \right) = \dot{F}, \\ F being a scalar function depending on \xi, \eta, \zeta and t, having as trivial case: F = const. In the Cartesian coordinates: \\ \end{array}$$

 $h_{\xi} = h_x = 1$ ;  $h_{\eta} = h_y = 1$ ;  $h_{\zeta} = h_z = 1$ , so that the integrability condition becomes  $\partial \dot{x} / \partial x + \partial \dot{y} / \partial y + \partial \dot{z} / \partial z = \nabla V = -\dot{F}$ .

Further two related rare cases of integrability for the continuity equation were analyzed. In all these cases the integrating procedure acts on either first or second term in the left-hand side of complete equation in vector form. Both cases (1 & 2) have a pair of subcases.

So, for the case 1 there are the subcases:

1.a. considering the usual local form:  $\partial \rho / \partial t + \nabla (\rho \mathbf{V}) = 0$ and setting:  $\rho = \nabla \mathbf{P}$ , and with  $\rho \mathbf{V} \neq \nabla \times \Psi_c$  ( $\rho \mathbf{V}$  is not a solenoidal vector, in order to keep the complete equation); 1.b. using the equivalent local form as in the beginning and setting:  $d(\ln \rho)/dt = \nabla \mathbf{Q}$ , and with  $\mathbf{V} \neq \nabla \times \Psi$  ( $\mathbf{V}$  is not a solenoidal vector, with the same reason as at 1.a). Here  $\mathbf{P}(\xi, \eta, \zeta, t) \neq \nabla \times \mathbf{W}$  is the Greek "Rho" vector and  $\mathbf{Q}(\xi, \eta, \zeta, t) \neq \nabla \times \mathbf{W}$  also, respectively (in order to assure the existence of functions  $\rho$  and  $d(\ln \rho)/dt$ , respectively,  $\mathbf{W}$  being a certain vector field). So both  $\mathbf{P}$  and  $\mathbf{Q}$  are not solenoidal vectors. Thus we get further: 1.a.  $\nabla (\partial \mathbf{P}/\partial t + \nabla \mathbf{P} \cdot \mathbf{V}) = 0$ , and 1.b.  $\nabla (\mathbf{Q} + \mathbf{V}) = 0$ . Hence the following first integrals are obtained:

1.a.  $\partial \mathbf{P}/\partial t + \nabla \mathbf{P} \cdot \mathbf{V} = \nabla \times \mathbf{Y}$ , with  $\mathbf{Y} \neq \nabla F_1$ , and 1.b.  $\mathbf{Q} + \mathbf{V} = \nabla \times \mathbf{Z}$ , with  $\mathbf{Z} \neq \nabla F_1$  also, ( $F_1$  being a certain scalar field). So both  $\mathbf{Y}$  and  $\mathbf{Z}$  are not conservative vectors, and thus the vectors ( $\partial \mathbf{P}/\partial t + \nabla \mathbf{P} \cdot \mathbf{V}$ ) and ( $\mathbf{Q} + \mathbf{V}$ ) are both solenoidal. One can see that among the main working hypotheses considered in this section were: 1.a.  $\rho \mathbf{V} \neq \nabla \times \Psi_c$ , and respectively 1.b.  $\mathbf{V} \neq \nabla \times \Psi$ , in both cases meaning that the mass flux density  $\mathbf{j} = \rho \mathbf{V}$  and the velocity  $\mathbf{V}$  respectively are no longer derived from the new introduced 3-D "stream function" vector  $\Psi_c$  (for a compressible fluid flow) and the similar 3-D true stream function vector  $\Psi$  (for a compressible fluid flow also), and all the results obtained in the previous sections cease to be valid for the complete equation.

The subcases 1.a and 1.b correspond to an unsteady flow of a special anisotropic fluid.

A second pair of subcases is:

2.a. considering the usual local form:  $\partial \rho / \partial t + \nabla (\rho \mathbf{V}) = 0$  and setting:  $\rho \mathbf{V} = \partial \mathbf{T} / \partial t$ , where  $\mathbf{T}(\xi, \eta, \zeta, t)$  is a special vector; 2.b. using the equivalent local form as in the beginning and setting:  $\nabla V = dU/dt$ , where U( $\xi$ ,  $\eta$ ,  $\zeta$ , t) is a special scalar function. We get further the following equations: 2.a.  $\partial(\rho + \nabla T)/\partial t = 0$ , and 2.b.  $d(\ln \rho + U)/dt = 0$ . In order to preserve the complete continuity equation in its both initial local forms (2.a) and (2.b), a special condition must be imposed on both functions T and U: 2.a.  $\mathbf{T} \neq \nabla \times \mathbf{W}$  and 2.b.  $dU/dt \neq 0$  also (W being a certain vector field). Hence the following first integrals: 2.a.  $\rho + \nabla \mathbf{T} = \mathbf{G}(\xi, \eta, \zeta) + \mathbf{C}_0$ , (G being a scalar field and C<sub>0</sub> a scalar constant), meaning a steady general behavior, and 2.b.  $\ln \rho + U = C_1$ , (C<sub>1</sub> being another scalar constant). We will determine for the subcase (2.b) the dependence  $\rho \leftrightarrow V$ , for simplicity reasons using Cartesian coordinates

$$U = \int \nabla V dt = \int \frac{\partial V_x}{\partial x} \frac{dt}{dx} dx + \int \frac{\partial V_y}{\partial y} \frac{dt}{dy} dy + \int \frac{\partial V_z}{\partial z} \frac{dt}{dz} dz$$
$$= \int \frac{1}{V_x} \frac{\partial V_x}{\partial x} dx + \int \frac{1}{V_y} \frac{\partial V_y}{\partial y} dy + \int \frac{1}{V_z} \frac{\partial V_z}{\partial z} dz$$
$$= \int \frac{\partial (\ln V_x)}{\partial x} dx + \int \frac{\partial (\ln V_y)}{\partial y} dy + \int \frac{\partial (\ln V_z)}{\partial z} dz; (F = -U)$$
-solution of equation  $\nabla V = -\dot{F}$ , and then  $\rho/\rho_i = e^{C_1 - U}$ .

13 The vorticity equation and its first integral; introducing the 2-D vorticity quasipotential and Selescu's roto-viscous vector; analogy to "magnetic induction" equation This equation describes the time variation of the vorticity intensity  $\Omega$  in a viscous Newtonian incompressible fluid flow. It can be derived from the motion equation (Navier–Stokes – see subsection 1.7 in [1], [3]), applying the curl and taking into account that the mass force density is conservative  $\mathbf{f} = -\nabla(\mathbf{gz})$  and the static pressure  $\mathbf{p} = f(\rho \text{ only})$ :  $\partial \Omega / \partial t = \nabla \times (\mathbf{V} \times \Omega) + \nu \Delta \Omega$ ; (11)

 $v = \mu_d / \rho$  is the kinematic viscosity of the fluid medium – a constant quantity. The first term in the right-hand side is a convective one, while the second term is a diffusive one. So far, first integrals were found for a single term in the righthand side only. Taking into account that  $\Omega$  is a solenoidal field ( $\nabla \Omega = 0$ ), and that  $\Delta \Omega = \nabla(\nabla \Omega) - \nabla \times (\nabla \times \Omega)$ , one gets:  $\Delta \Omega = -\nabla \times (\nabla \times \Omega)$ , and so:  $\partial \Omega / \partial t = \nabla \times [\nabla \times \Omega - \nabla \nabla \Omega]$  $v(\nabla \times \Omega)$ ] or (with  $\Theta = \nabla \times \Omega$ ):  $\partial \Omega / \partial t = \nabla \times (\nabla \times \Omega - v\Theta)$ . On any  $(\Omega, \Theta)_i$  surface we can introduce a new smart triorthogonal intrinsic coordinate system  $O\xi_1\eta_1\zeta_1$  tied to these surfaces (or  $O\lambda_1\mu_1\nu_1$ ; with  $\lambda_1, \mu_1, \nu_1$  – orthogonal arcs lengths;  $\lambda_1$ ,  $\mu_1$  – contained in the local tangent plane and  $v_1$  directed along the normal). Let us introduce the scalar function  $X_i(M) = X_i(\xi_1, \eta_1, \zeta_{10i})$  – the Greek "Chi", a 2-D "quasi-potential", whose partial derivatives along the directions of the elementary orthogonal arcs  $d\lambda_1 = h_{\xi_1} d\xi_1$ and  $d\mu_1 = h_{\eta_1} d\eta_1$  on the "i"  $(\Omega, \Theta)$  surface  $(\zeta_1 = \zeta_{10i})$  are just the components  $\Omega_{\text{Eli}}$  and  $\Omega_{\text{nli}}$  of the vorticity  $\Omega$  vector:

$$\begin{split} \Omega_{\xi li} &= \frac{1}{h_{\xi l}} \frac{\partial X_i}{\partial \xi_1} = \frac{\partial X_i}{\partial \lambda_1} \quad ; \quad \Omega_{\eta li} = \frac{1}{h_{\eta}} \frac{\partial X_i}{\partial \eta_1} = \frac{\partial X_i}{\partial \mu_1} \quad ; \\ \Omega_{\zeta li} &= \frac{1}{h_{\zeta l}} \frac{\partial X_i}{\partial \zeta_1} = \frac{\partial X_i}{\partial \nu_1} = 0 \quad , \quad \Omega_i = \nabla X_i \quad . \end{split}$$

1) A trivial case of first integrability is given by:  $\mathbf{V} \times \mathbf{\Omega} - v(\nabla \times \mathbf{\Omega}) = \nabla \mathbf{F}$  (F being a scalar function and  $\nabla \mathbf{F}$  a pure gradient), getting for the right-hand side of the vorticity equation the form:  $\nabla \times \nabla \mathbf{F} = 0$  – a particular case ( $\partial \mathbf{\Omega} / \partial t = 0$  – a steady vortex field). Performing a scalar multiplication of this equation by a virtual elementary displacement d**R** normal to the vector  $\mathbf{v}_i \cdot \mathbf{v} = \mathbf{V} \times \mathbf{\Omega} - v(\nabla \times \mathbf{\Omega})$  – the viscous incompressible vector, one gets:  $[\mathbf{V} \times \mathbf{\Omega} - v(\nabla \times \mathbf{\Omega})] \cdot d\mathbf{R} = 0$ , therefore a "zero-work" condition for the vector  $\mathbf{v}_i \cdot \mathbf{v} \cdot (\mathbf{v}_i \cdot \mathbf{v} \cdot d\mathbf{R} = 0)$ , this meaning a surface – envelope sheet of the local planes normal to the **v**<sub>i</sub>.**v**. vector, one obtains:  $\nabla F \cdot d\mathbf{R} = dF = 0$ , and the first integral for the steady case: F = const.2) An important case of integrability, after performing a scalar multiplication of the equation by a virtual  $d\mathbf{R}$ , is that one consisting in simultaneously satisfying two conditions, to eliminate the non-conservative terms: 2a) d**R** – coplanar with both  $\Omega$  and  $\Theta$ ; this may be expressed as:  $d\mathbf{R} = c_3 \mathbf{\Omega} + c_4 \Theta$  (with:  $d\mathbf{R} \perp (\Theta \times \mathbf{\Omega})$ ); 2b) d**R** – normal to the vector  $\mathbf{W}_i = \nabla \times \mathbf{v}_i \cdot \mathbf{v}_i$ .  $= \nabla \times [\mathbf{V} \times \mathbf{\Omega} - \mathbf{v}(\nabla \times \mathbf{\Omega})]$ , or in the form of a zeroscalar product:  $\{\nabla \times [\mathbf{V} \times \mathbf{\Omega} - \mathbf{v}(\nabla \times \mathbf{\Omega})]\} \cdot d\mathbf{R} = 0$ , therefore a "zero-work" condition for  $W_i$  ( $W_i \cdot d\mathbf{R} = 0$ ). The first condition leads to a  $(\Omega, \Theta)$  surface – envelope sheet of the  $(\Omega, \Theta)$  local planes, having  $(\Theta \times \Omega) \cdot d\mathbf{R} = 0$ . Like in subsection 1.6 in [1], introducing a 2-D vorticity "quasi-potential"  $X_i$ , and writing  $\Omega_i = \nabla X_i$  (where, this time  $X_i$  is a scalar function depending not only on  $\xi_1$ ,  $\eta_1$  or  $\lambda_1, \mu_1$  but on t also), if the virtual elementary d**R** vector is contained in the  $(\Omega, \Theta)$  tangent plane, integrating the term  $\partial (\nabla X_i) / \partial t \cdot d\mathbf{R}$ , one gets the term  $\partial X_i / \partial t$  (due to Schwarz' theorem – similarly to the case in subsection 1.6 in [1]). The second condition means a surface – envelope sheet of the local planes normal to W<sub>i</sub> vector. Solving this system one gets some space curves - intersection lines of the envelope sheets above - along which the searched for dR must be directed, for the vorticity equation to admit a first integral. These lines are the unique solution of the system given by the conditions 2a and 2b:

$$\begin{split} d {\bm R} &\in (\Omega, \Theta) \text{ and also } d {\bm R} \perp W_i \text{ , this resulting in:} \\ d {\bm R} &\in ((\Omega, \Theta) \cap (\text{plane normal to } W_i)) \parallel \& -\text{ the } \& \\ (\text{ampersand}) \text{ vector lines, } d {\bm R} \text{ being therefore parallel to} \\ \text{the vector: } \& = (\Theta \times \Omega) \times [\nabla \times (V \times \Omega - v\Theta)] \text{ , with:} \\ \& -\text{ Selescu's vector, representing } a \text{ new physical} \\ \textit{quantity (incompressible roto-viscous), and giving for the} \\ \text{most general case of vorticity equation the first integral:} \\ \partial X_i / \partial t = C_{0i}(t), \end{split}$$

for the entire equation (the local derivative  $\partial \Omega / \partial t$  and both convective  $\nabla \times (\mathbf{V} \times \mathbf{\Omega})$  and diffusive  $v \Delta \mathbf{\Omega}$  term). 3) Another particular case of first integrability, also after performing a scalar multiplication of the equation by a virtual elementary displacement dR vector, is that one consisting in simultaneously satisfying other two conditions, to eliminate the non-conservative terms:  $(3a) \equiv 2a) d\mathbf{R}$  – coplanar with both  $\boldsymbol{\Omega}$  and  $\boldsymbol{\Theta}$ ; this may be expressed as:  $d\mathbf{R} = c_3 \mathbf{\Omega} + c_4 \Theta$  (with:  $d\mathbf{R} \perp (\Theta \times \Omega)$ ); 3b)  $\mathbf{W}_{i} = \nabla F_{1} + \mathbf{C}_{1}(t)$  or:  $\nabla \times [\mathbf{V} \times \mathbf{\Omega} - v(\nabla \times \mathbf{\Omega})] = \nabla F_{1}$ +  $C_1(t)$  (F<sub>1</sub> being a scalar function and  $\nabla F_1$  a gradient). getting for the differential equation of the vorticity vector:  $\partial \Omega / \partial t = \nabla F_1 + C_1(t)$ , or performing a scalar multiplication by the virtual displacement d**R** given by the condition 3a):  $[\partial(\nabla X_i)/\partial t] \cdot d\mathbf{R} = [\nabla F_1 + C_1(t)] \cdot d\mathbf{R} \quad (C_1 - a \text{ vector}),$ or (due to Schwarz' theorem applied to the left-hand side):  $d(\partial X_i/\partial t) = dF_1 + C_1(t)dR$ , so yielding the first integral:

#### $\partial \mathbf{X}_i / \partial \mathbf{t} = \mathbf{F}_1 + \mathbf{C}_1(\mathbf{t})\mathbf{R} + \mathbf{C}_2(\mathbf{t})$

(again for the entire equation); the scalar function  $F_1$  must be determined from the condition 3b), this giving a special relation between V and  $\Omega$ . C<sub>1</sub> and C<sub>2</sub> are arbitrary functions of t. By analogy the "magnetic induction" equation in MHD (with the analogy laws:  $\mathbf{H} \leftrightarrow \mathbf{\Omega}$ ;  $v_m \leftrightarrow v$ ) was studied:  $\partial \mathbf{H} / \partial \mathbf{t} = \nabla \times (\mathbf{V} \times \mathbf{H}) + \mathbf{v}_{\mathrm{m}} \Delta \mathbf{H}$ (12)(the equation of the magnetic field intensity; see for reference Eq. (11)) which was treated in [26] (subsection 2.9), where: **H** is the intensity of the local magnetic field, making no distinction between H and the magnetic induction B (in the Gaussian system of units), since for all electroconducting fluids the magnetic permeability  $(\mu)$  is approximately equal to 1 (see for reference [27] - [30]); V is the mean instantaneous velocity of translation of the ionized fluid (plasma) particles (atoms, ions etc.) contained in a local small volume element (see subsection 2.1 in [26]):  $\mathbf{V} = (\rho_a \mathbf{V}_a + \rho_+ \mathbf{V}_+ + \rho_- \mathbf{V}_-)/\rho$ , where:  $\mathbf{V}_a$ ,  $\mathbf{V}_+$ ,  $\mathbf{V}_-$  are the velocities of the components;  $\rho_a$ ,  $\rho_+$ ,  $\rho_-$  are the densities of the components:  $\rho_a = n_a m_a$ ;  $\rho_+ = n_+ m_+$ ;  $\rho_- = n_- m_-$ ;  $\rho = n_a m_a + n_+ m_+ + n_- m_- = \rho_a + \rho_+ + \rho_- - \text{plasma density}$ (analogously to the case of a mixture of components);  $m_a$  is the mass of a neutral atom ( $m_a = m_+ + m_-$ );  $m_+$  is the mass of the positive ion (a single species);  $m_{-}$  is the mass of the negative ion or of electron;  $n_a$  is plasma concentration in neutral atoms, and respectively (for a threecomponent neutral plasma):  $n_+$ ,  $n_-$  are the concentrations in positive and negative particles (a single species of cations and anions; in the case of a quasi-neutral plasma:  $n_+ \approx n_-$ ), all according to a simplified model proposed by the author;  $v_m$  is the magnetic viscosity of the electroconducting fluid. Analogously to  $X_i(M) = X_i(\xi_1, \eta_1, \zeta_{10i})$ , with  $\Omega_i = \nabla X_i$ , we can introduce a 2-D magnetic "quasi-potential"  $\Xi_i(M)$  $= \Xi_i (\xi_2, \eta_2, \zeta_{20i})$ , with  $\mathbf{H}_i = \nabla \Xi_i$ , whose partial derivatives along the directions of the elementary orthogonal arcs  $d\lambda_2$ =  $h_{\xi_2}d\xi_1$  and  $d\mu_2 = h_{\eta_2}d\eta_2$  on any "i" (**H**,  $\nabla \times$  **H**) surface ( $\zeta_2$ =  $\zeta_{20i}$ ) are just the components  $H_{\xi_{2i}}$  and  $H_{\eta_{2i}}$  of the vector **H**. A new physical quantity (similar to the incompressible rotoviscous vector &) was introduced – Selescu's magnetic vector:  $\&_m = [(\nabla \times \mathbf{H}) \times \mathbf{H}] \times \{\nabla \times [\nabla \times \mathbf{H} - \nu_m(\nabla \times \mathbf{H})]\}.$ The low-frequency Ampère's law neglects displacement current, being:  $\mathbf{j} = k\nabla \times \mathbf{H}$ , with  $k = c/4\pi$ ; c is the light speed in a vacuum; i is the density of conduction current. So we get:  $\&_{\rm m} = (\mathbf{j}/k \times \mathbf{H}) \times [\nabla \times (\mathbf{V} \times \mathbf{H} - v_{\rm m}\mathbf{j}/k)]$ .

So far, to the best of the author's knowledge, *only two particular cases of first integrability* are known for Eq. (12): 1. *ideal conducting plasma* (with zero-resistivity and thus zero-magnetic viscosity, missing the diffusion term  $v_m\Delta H$ ), leading to the concept of magnetic field "frozen" in plasma (given by an equation with the same form as Helmholtz' one [19] for the vorticity in the ideal fluid in hydrodynamics) *and* 2. *non-moving real plasma* (with finite conductivity and thus finite magnetic viscosity, now missing the convection term  $\nabla \times (\mathbf{V} \times \mathbf{H})$ ), getting a standard diffusion equation.

#### 14 Conclusions and remarks

The contribution of this work to the state of the art in fluid mechanics (especially aerothermodynamics and MHD), as well as in electromagnetic theory, consists in finding new first integrability cases and first integrals for: 1. *the continuity equation* (for the steady flow of an inviscid compressible fluid), obtaining some special general forms of the flow rate equation;

2. *the system of equations (motion and continuity)* for the steady flow of an inviscid compressible fluid;

3. the general case and two particular cases of complete continuity equation (for the unsteady flow of an inviscid compressible fluid), each of them with two subcases; 4. the complete vorticity equation (with both convection and diffusion terms) for a viscous incompressible fluid flow; 5. the complete "magnetic induction" equation in MHD (with both terms) for the flow of a magnetic viscous electroconducting fluid (plasma included), by analogy to previous one. Three smart intrinsic curvilinear coordinate systems were used:  $O\xi\eta\zeta$  (or  $O\lambda\mu\nu$ ), for the continuity equation,  $O\xi_1\eta_1\zeta_1$  (or  $O\lambda_1\mu_1v_1$ ), for the vorticity equation, and  $O\xi_2\eta_2\zeta_2$  (or  $O\lambda_2\mu_2v_2$ ), for the "magnetic induction" one. In this work a rich nomenclature was introduced in fluid mechanics (for both inviscid and viscous fluids), as well as in viscous MHD and electromagnetic theory: - the 2-D "quasi-stream"  $F_i$  function on the 3-D (V,  $\Omega$ ) "i" surfaces, for the continuity equation;

- the surfaces of *iso*-normal mass flux *density*  $|\mathbf{jn}| = \rho |\mathbf{Vn}|$  $= (\rho |V_n| =)$  const. (on which the continuity equation of the steady flow of a compressible fluid in a thick stream tube admits the same first integral as for the fluid flow in a thin stream tube, and whose envelope sheets are just the sections of uniform flow, on which  $\rho = \text{const.}$  and  $|V_n|$ = const. as well, if they exist), for the flow rate equation; - the 3-D stream function vector  $\Psi_{c}$  (Selescu, defined by:  $\nabla \times \Psi_{c} = \rho \mathbf{V}$ , with  $\Psi_{c} \neq \nabla G$ ), allowing a new global form for the continuity equation of a compressible fluid steady flow (a stream function for the mass flux density  $\rho$ **V**), using the circulation of  $\Psi_c$  along a closed curve  $l_i$ bounding a certain cross section  $S_i$  of a thick stream tube and applying the classical Kelvin-Stokes theorem (a reciprocal one), similarly to Helmholtz' 2nd theorem about vortices in an inviscid fluid, using the circulation of V along any closed curve surrounding a vortex tube:  $(\Gamma_{\Psi_c})_{out} = -(\Gamma_{\Psi_c})_{in} = \text{const.}; \ (\Gamma_{\Psi_c})_{in, out} = \int_{I_{in}, I_{out}} \Psi_c \cdot d\mathbf{R};$ - three 2-D "quasi-potential" scalar functions: a velocity one,  $\Phi_i(M)$ , on the 3-D (V,  $\Omega$ ) surfaces (V<sub>i</sub> =  $\nabla \Phi_i$ ); a vorticity one,  $X_i(M)$ , on the 3-D ( $\Omega, \Theta$ ) surfaces ( $\Omega_i = \nabla X_i$ ); and a magnetic one,  $\Xi_i(M)$ , on the 3-D (**H**, **j**) surfaces (**H**<sub>i</sub> =  $\nabla \Xi_i$ ); - the virtual "zero-work" surfaces for the vectors  $\mathbf{v_{i}}$ .  $\mathbf{V} \times \mathbf{\Omega} - v(\nabla \times \mathbf{\Omega})$  and  $\mathbf{W}_i = \nabla \times \mathbf{v}_i \cdot \mathbf{v}_i$  (in vorticity equation). Two new vector physical quantities were introduced:  $\& = (\Theta \times \Omega) \times [\nabla \times (\nabla \times \Omega - \nu \Theta)] - \text{Selescu's incompres-}$ sible roto-viscous one, for a certain 3-D flow, and by analogy,

in MHD:  $\&_m = (j/k \times H) \times [\nabla \times (V \times H - v_m j/k)]$  – Selescu's *magnetic* one, to find lines along which we get first integrals. *The interdependence* between the 2-D quasi-potential  $\Phi_i$  and the 2-D quasi-stream function  $\Psi_{ci}$  was also established. *Any continuity equation* (in electromagnetic theory and MHD) can be treated in a manner similar to this one. In plasma-MHD the equation has the same form as in fluid mechanics, with a slightly different nomenclature (section 13). Finally we can cite a remarkable sentence from [31]: "Not always the simplest is the best explanation. The analogies sometimes work and sometimes not. Old or conventional theories may be misleading."

#### Note

This paper (the second in a series dedicated to the intrinsic analytic study of the basic equations in compressible fluid mechanics) is fully original, however having as starting point other ones with almost the same titles (see [32], [4]).

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